




The National Superconducting Cyclotron Laboratory

Michigan State University

U.S. flagship user facility for rare isotope research and education in nuclear science, astro-nuclear physics, accelerator physics, and societal applications

Betty Tsang for the  Collaboration
JCNP2015, Osaka,
November 8, 2015



Nuclear Symmetry Energy: From Nucleus to Neutron Stars

USA

State and Capital

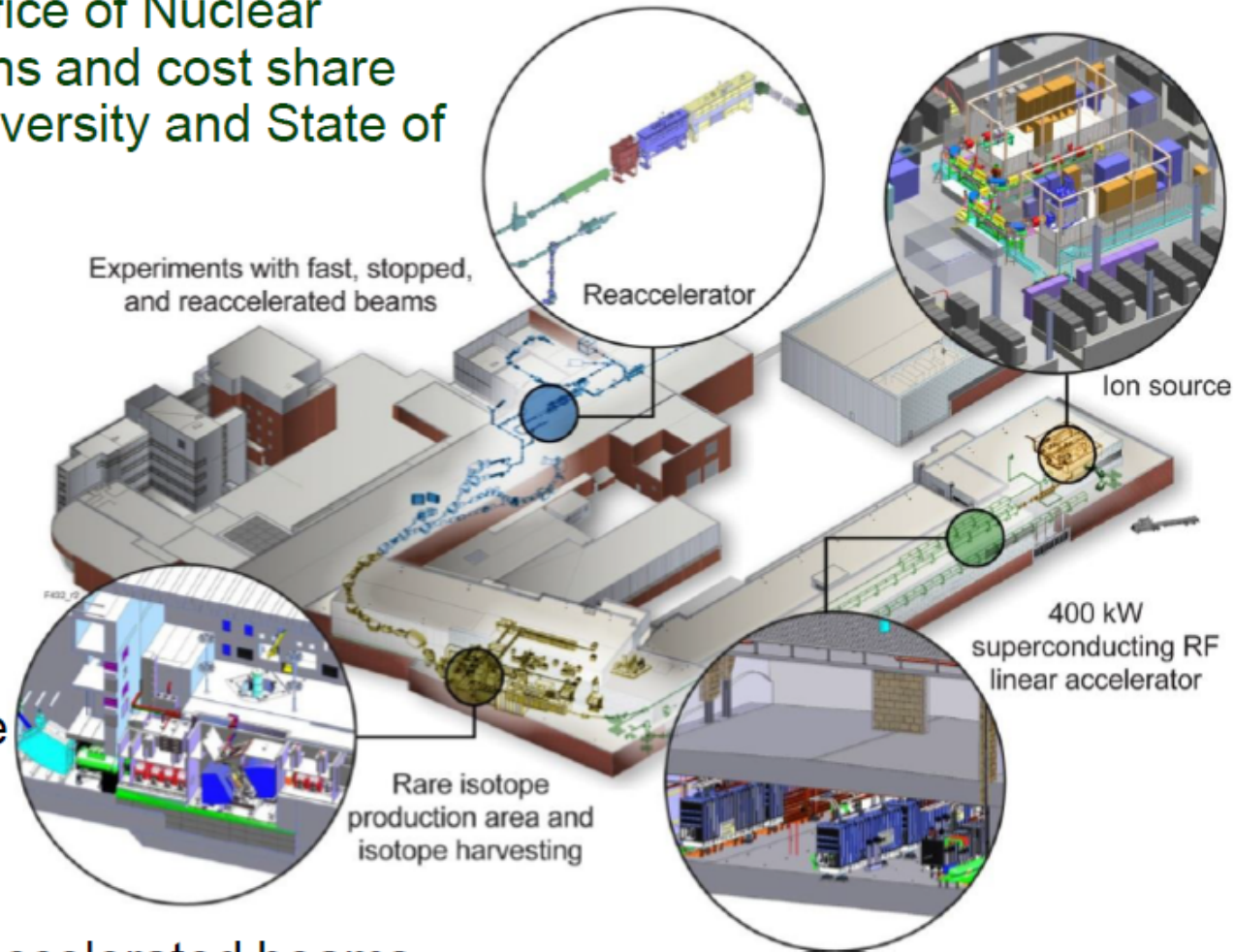




Michigan State University

From NSCL to Facility for Rare Isotope Beams (FRIB)

- Funded by DOE–SC Office of Nuclear Physics with contributions and cost share from Michigan State University and State of Michigan
- Managing to early completion in Dec 2020
- Key feature is 400 kW beam power for all ions (5×10^{13} $^{238}\text{U/s}$)
- Separation of isotopes in-flight
 - Fast development time for any isotope
 - Suited for all elements and short half-lives
 - Fast, stopped, and reaccelerated beams

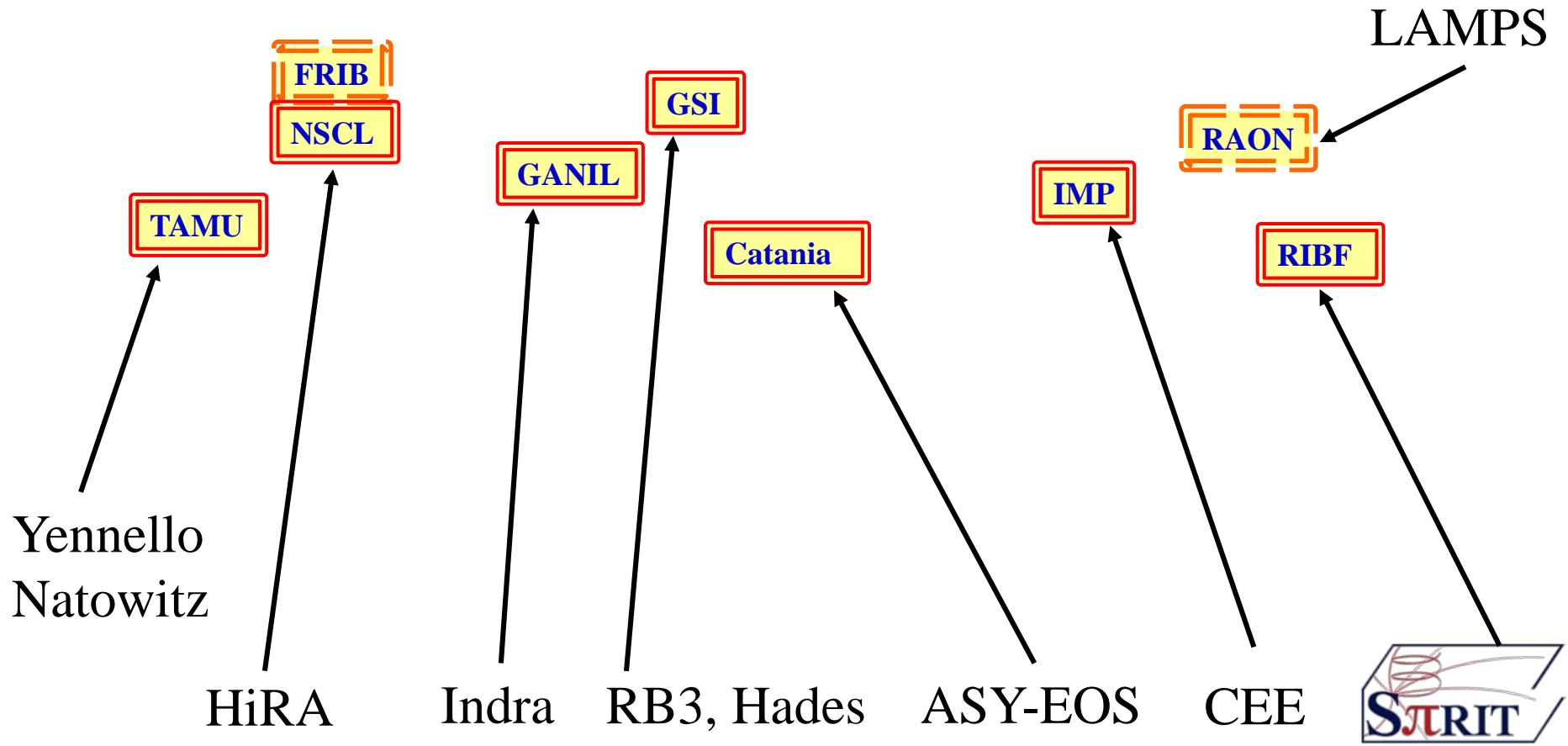
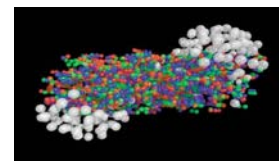




FRIB Construction Progress



Symmetry Energy Project



International Collaboration is the key



JUSTIPEN

Japan-U.S. Theory Institute for Physics with Exotic Nuclei

- JUSTIPEN Home
- JUSTIPEN Visit Application Form
- Conferences, Schools, and Seminars
- Visitors
- Visitor Information
- Publications
- Collaborations
- Job Postings
- JUSTIPEN Policies
- Governing Board
- RNB Resources
- Photo Gallery
- Links
- News



About JUSTIPEN

Purpose:

Deliver an international venue for research on the physics of nuclei during an era of experimental investigations on rare isotopes.

Location:

RIKEN, at the new RIB Factory

US Participation:

Provide travel and local support for U.S. visits to JUSTIPEN

Synopsis:

The Japan-U.S. Theory Institute for Physics with Exotic Nuclei (JUSTIPEN) has been established in order to facilitate collaboration between U.S. and Japanese scientists whose main research is in the area of the physics of nuclei. JUSTIPEN is located at RIKEN RIB Experimental Facility in Wako, near Tokyo, Japan. Participation in JUSTIPEN is in the form of travel grants and subsistence grants to those individuals who are interested in collaborating with Japanese scientists. JUSTIPEN's purview includes the area of physics of or with exotic nuclei, including nuclear structure and reaction theory, nuclear astrophysics, and tests of the standard model using exotic nuclei. While JUSTIPEN primarily focuses

JUSTIPEN provide funds for US scientists to travel to RIBF... to collaborate with Japanese scientists



2011-FUSTIPEN



2006-2014



Time to change “Theory” to “Experiment)” (2009)



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JUSTIPEN provide funds for US scientists to travel to RIBF... to collaborate with Japanese scientists



- Home
- JUSEIPEN Organization
- JUSEIPEN Policies
- Grant Application Form
- Meetings
- JUSEIPEN related links

About JUSEIPEN

The Japan-US Experimental Institute for Physics with Exotic Nuclei (JUSEIPEN) has been established to help promote and develop strong US-Japan experimental collaborations using rare isotope beams.

JUSEIPEN can provide funds for U.S. scientist to travel to the RI Beam factory (RIBF) at the RIKEN Nishina Center in Wako near Tokyo, and to other institutes in support of JUSEIPEN's goals. JUSEIPEN can also provide a framework to organize and support workshops and meetings to help coordinate effort on the science goals and logistics for U.S. involvement in RIBF.

If you are a scientist working at a US institution interested in developing or joining a collaboration with Japanese colleagues to carry out experiments at the RI Beam factory, RIKEN, and you think JUSEIPEN can help, then please feel free contact the steering committee (via pfallon@lbl.gov). Information on applying for a travel funds is given under JUSEIPEN Policies.

JUSEIPEN provide funds for US scientists to travel to RIBF...

EN's activities is carried
tives from the U.S. and
the Office of Nuclear
The present cycle of
ugh FY2013.



CUSTIPEN

中美奇特核物理理论研究所

China-U.S. Theory Institute for Physics with Exotic Nuclei



The First FRIB-China Workshop on Physics of Nuclei and Hadrons

May 28-30, NSCL, Michigan State University

US-China-RIB.org

*connecting our common
research interests in
exotic nuclei*

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[Breaking
News](#)

[Science](#)

[Task
Force](#)

[Working
Groups](#)

[Gather-
ings](#)

[Documents](#)

[Links](#)

[Home](#)



Nuclear Symmetry Energy: From Nucleus to Neutron Stars

曾敏兒 -- Betty Tsang

Outline

1. Introduction : Different forms of EoS
2. How did we get here? Current constraints on density dependence of symmetry energy.
3. Where are we going? Future challenges and opportunities.
4. Research funded by DOE, NEXT and JUSEIPEN.
5. Summary

Nuclear Symmetry Energy: From Nucleus to Neutron Stars

Equation of State of nuclear matter

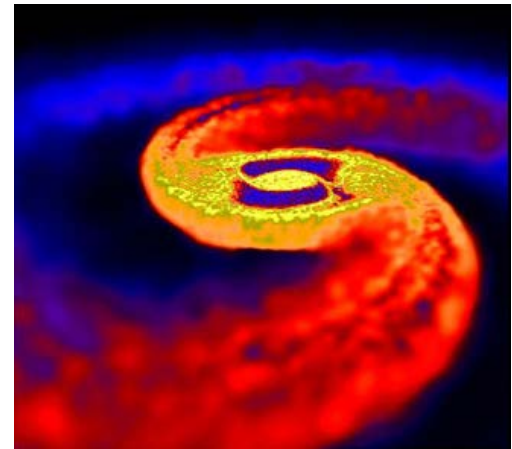
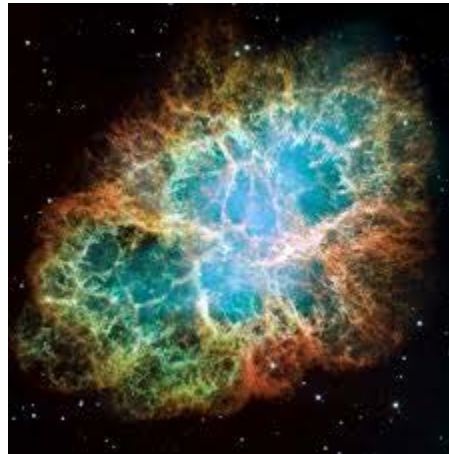
$$E/A(\rho, \delta) = E/A(\rho, 0) + \delta^2 \cdot S(\rho)$$

$$\delta = (\rho_n - \rho_p) / (\rho_n + \rho_p) = (N - Z) / A$$

Symmetry Energy of asymmetric matter



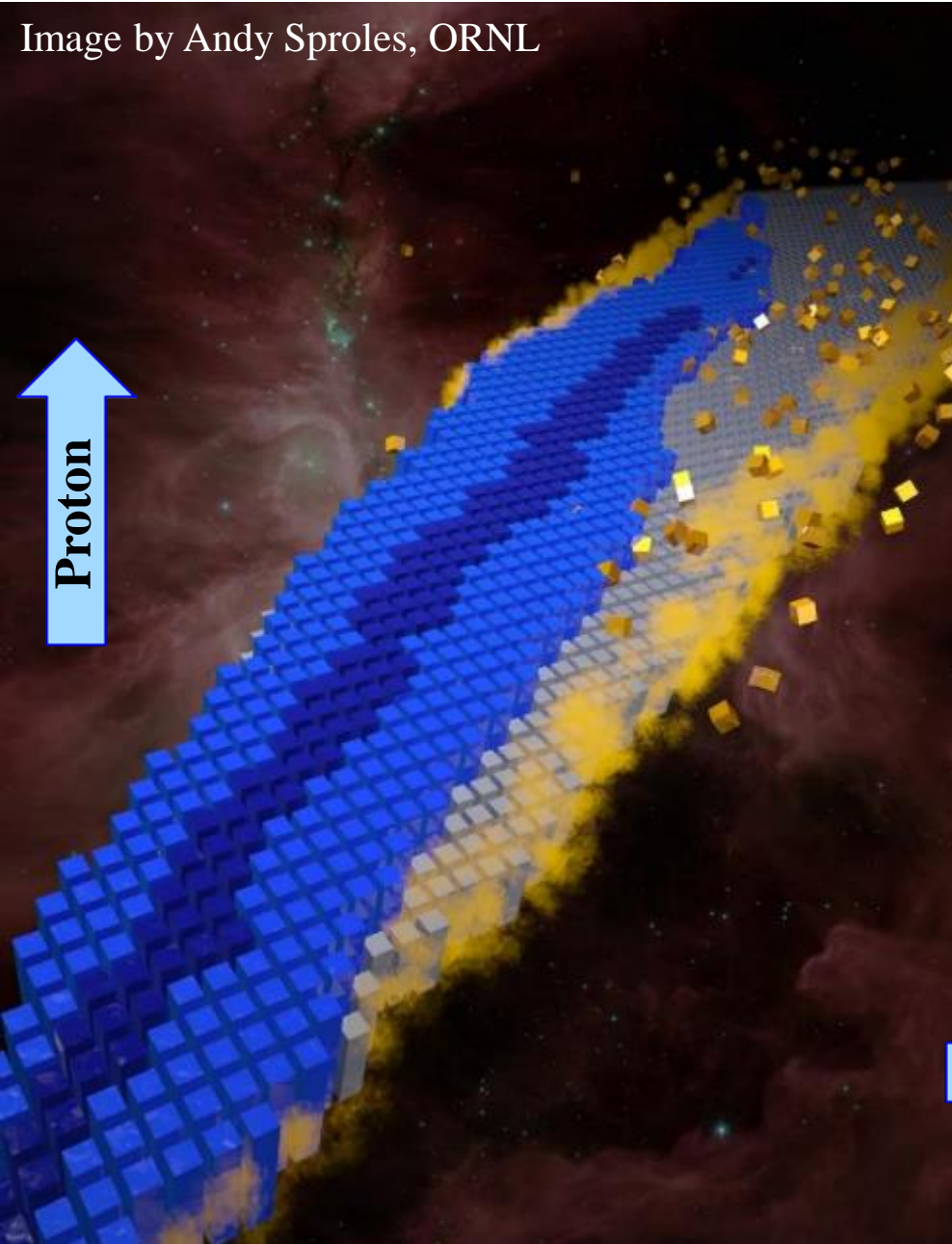
^{208}Pb



- To probe fundamental questions on the nature of nuclear matter especially the isospin asymmetric matter.
- To recreate and study astrophysical environments

Symmetry Energy

Image by Andy Sproles, ORNL



$$B = a_V A - a_S A^{2/3} - a_C \frac{Z(Z-1)}{A^{1/3}}$$

$$- a_{sym} \frac{(A - 2Z)^2}{A}$$

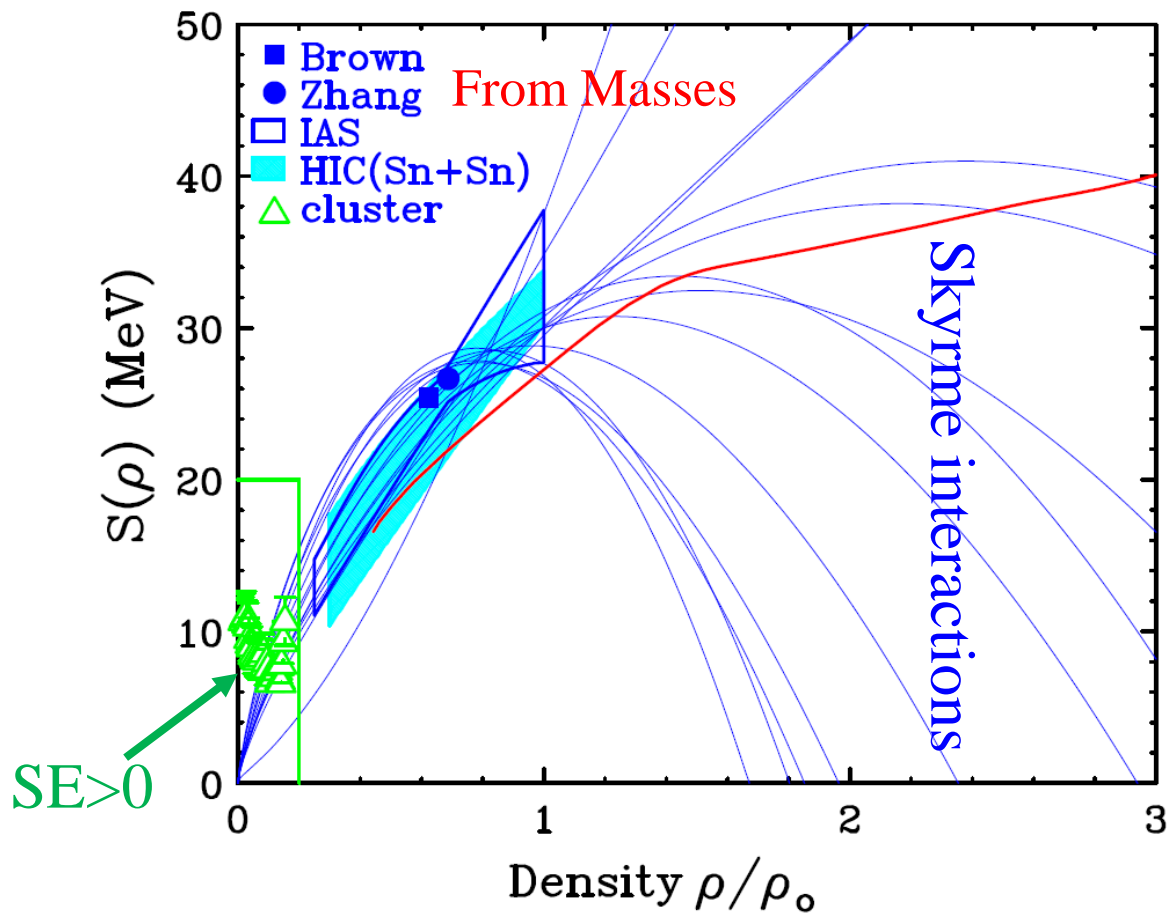


$$(a_{sym}^V A - a_{sym}^S A^{2/3}) \frac{(A - 2Z)^2}{A^2}$$

Inclusion of surface terms in symmetry



Hubble ST



Recommendation to
US Long Range Plan
from EOS working
group

- At $\rho \ll \rho_0$: Establish observables to study cluster effect and link to neutrinosphere physics.
- At $\rho \leq \rho_0$: Improve constraints from both structure and reaction experiments:
- At $\rho \approx 1.5 - 2\rho_0$: Determine symmetry energy and the momentum dependence of the isovector potential.

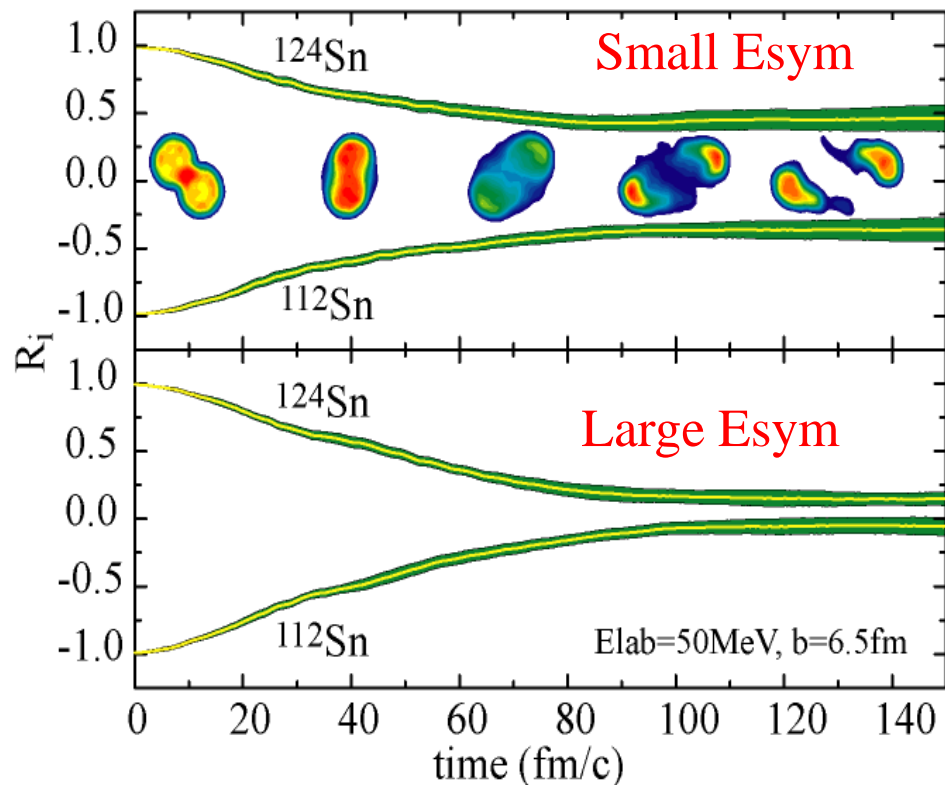
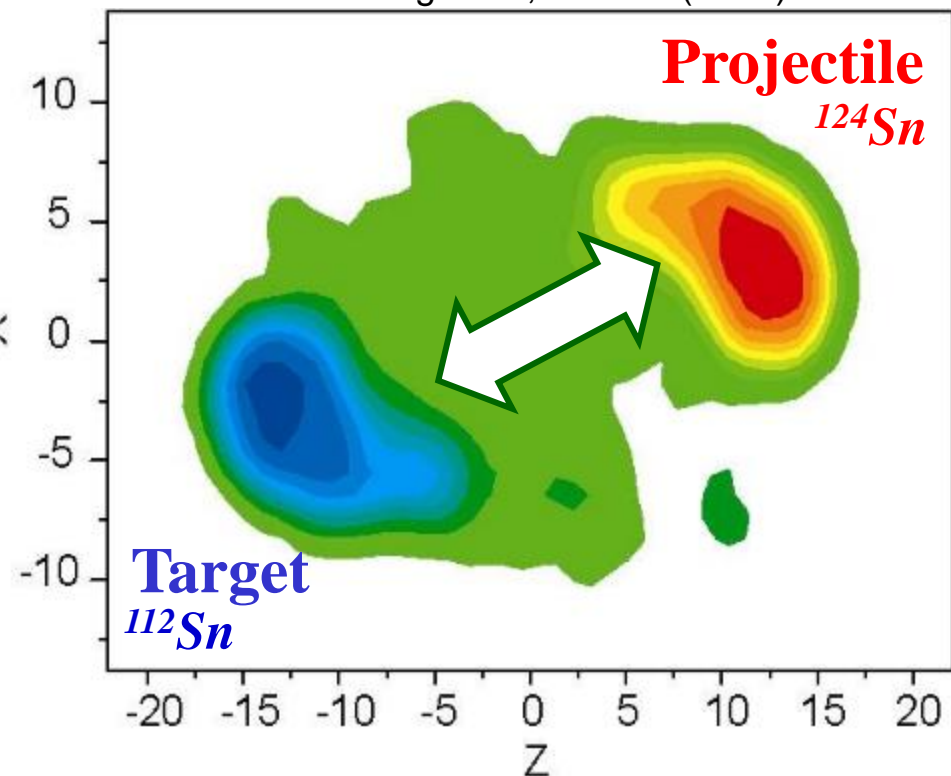
Isospin Diffusion observable to study E_{sym} with Heavy Ion Collisions

$$\delta = (N-Z)/A$$

Tsang et al., PRL 92 (2004) 062701

$$S(\rho) = 12.5(\rho/\rho_0)^{2/3} + C(\rho/\rho_0)^{\gamma_i}$$

Tsang, Shi et al., PRL92, 062701(2004)



Isospin Diffusion; low ρ , E_{beam}

$$R_i = 2 \frac{\delta_{AB} - (\delta_{AA} + \delta_{BB})/2}{\delta_{AA} - \delta_{BB}}$$

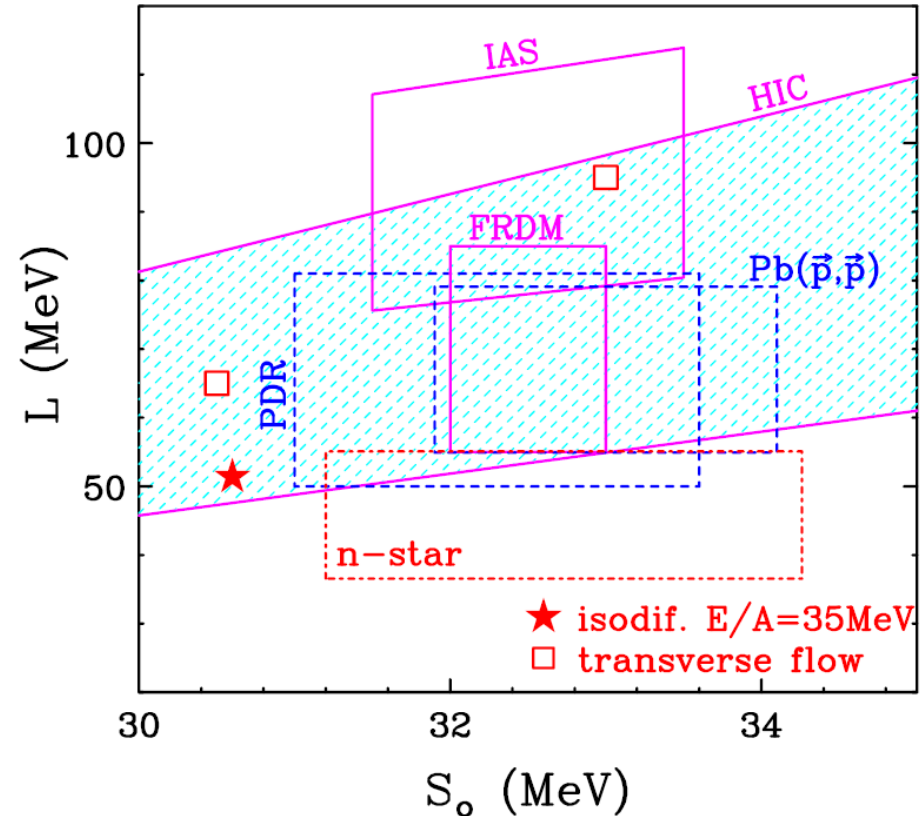
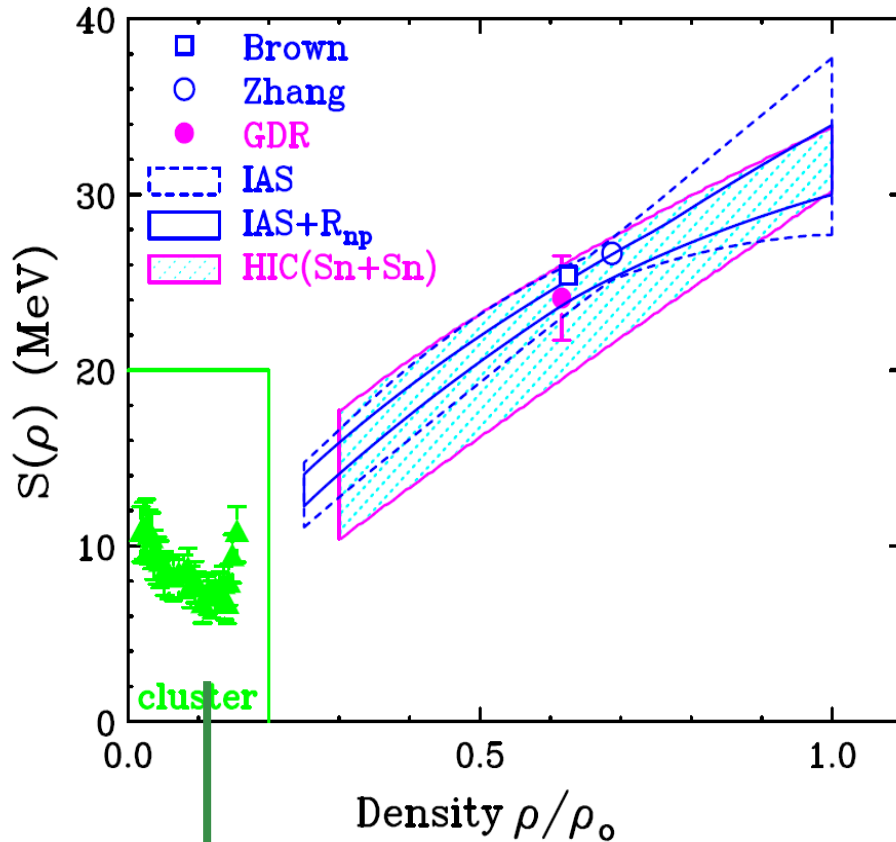
Bao-An Li et al., Phys. Rep. 464, 113 (2008)

Tsang, Zhang et al., PRL122, 122701(2009)

Status of Constraints from nuclear structure and reactions

$$S(\rho) = 12.5(\rho/\rho_0)^{2/3} + C(\rho/\rho_0)$$

$$S(\rho) = S_o + \frac{L}{3} \left(\frac{\rho_B - \rho_0}{\rho_0} \right) + \dots$$

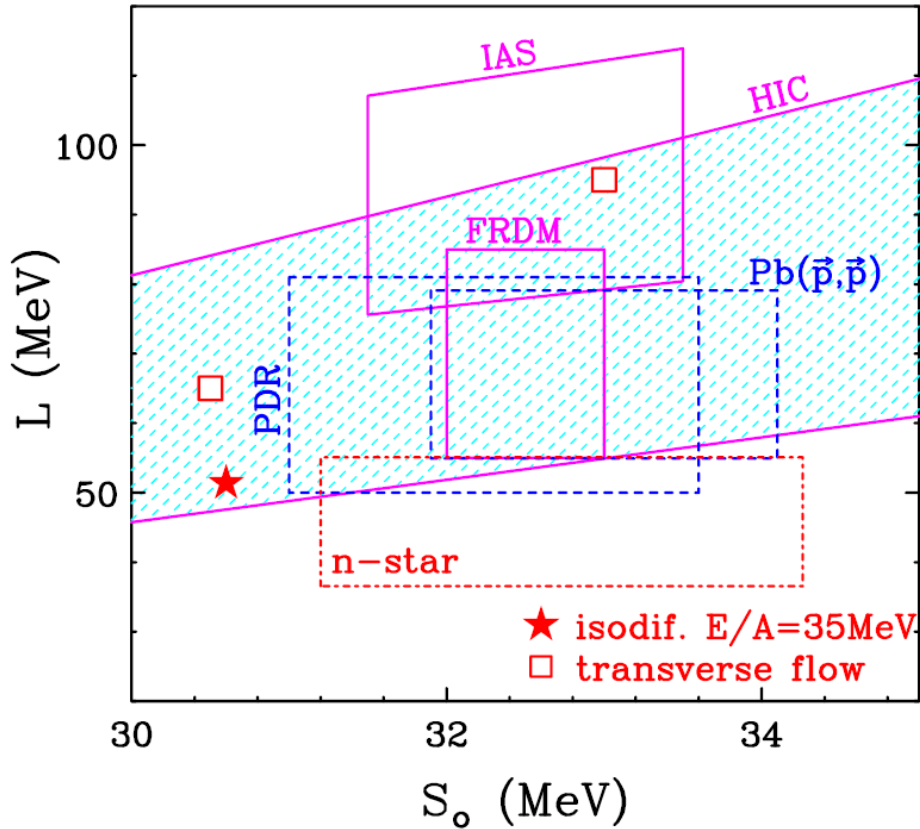


Consistent Constraints from nuclear structure and reactions with credible uncertainties

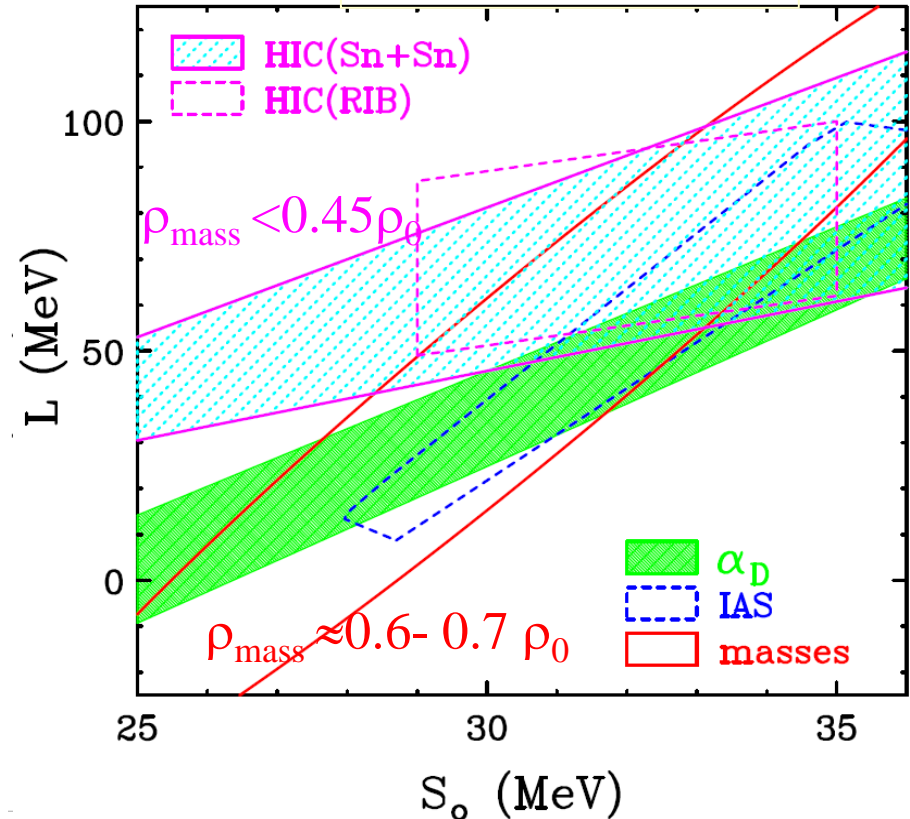
NuSYM13 & ICNT2013

$$S(\rho) = S_o + \frac{L}{3} \left(\frac{\rho_B - \rho_0}{\rho_0} \right) + \dots$$

Tsang et al, 86, 015803 (2012)



$$S(\rho) = 12.5(\rho/\rho_0)^{2/3} + C(\rho/\rho_0)$$



$$\rho_{\text{sens.}} = \rho_0 (1 - 3/M); \text{ M is slope}$$

Equation of State of Neutron Matter



Hubble ST

Neutron Star: balance of Gravity (pulls in) and Symmetry energy pressure (pushes out): Masses vs. Radii

$$\frac{dM}{dr} = 4\pi r^2 \mathcal{E}(r)$$

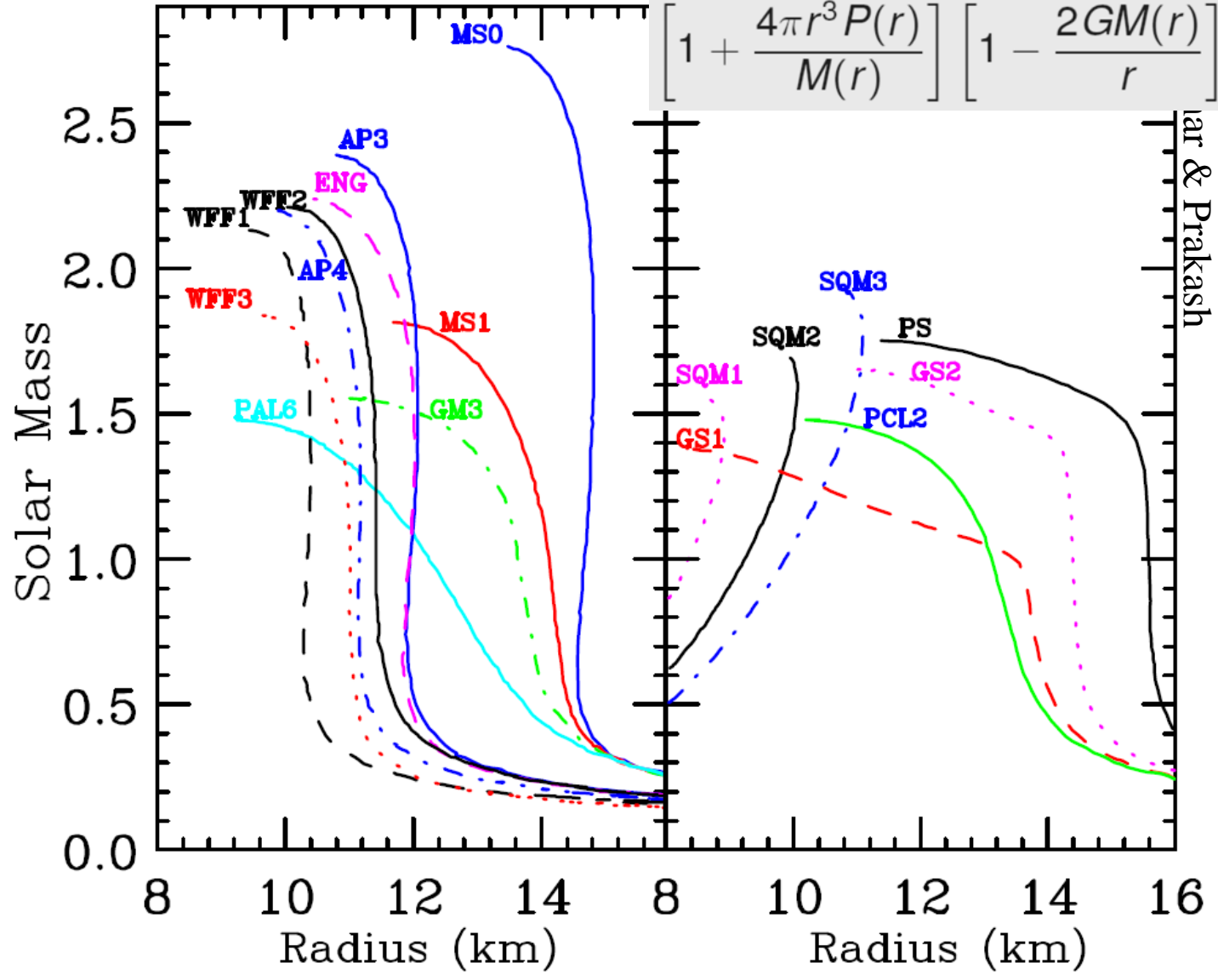
$$\frac{dP}{dr} = -G \frac{\mathcal{E}(r)M(r)}{r^2} \left[1 + \frac{P(r)}{\mathcal{E}(r)} \right] \left[1 + \frac{4\pi r^3 P(r)}{M(r)} \right] \left[1 - \frac{2GM(r)}{r} \right]^{-1}$$

EoS of pure neutron matter:
Symmetry Energy as function
of pressure (density)

$$\frac{dM}{dr} = 4\pi r^2 \mathcal{E}(r)$$

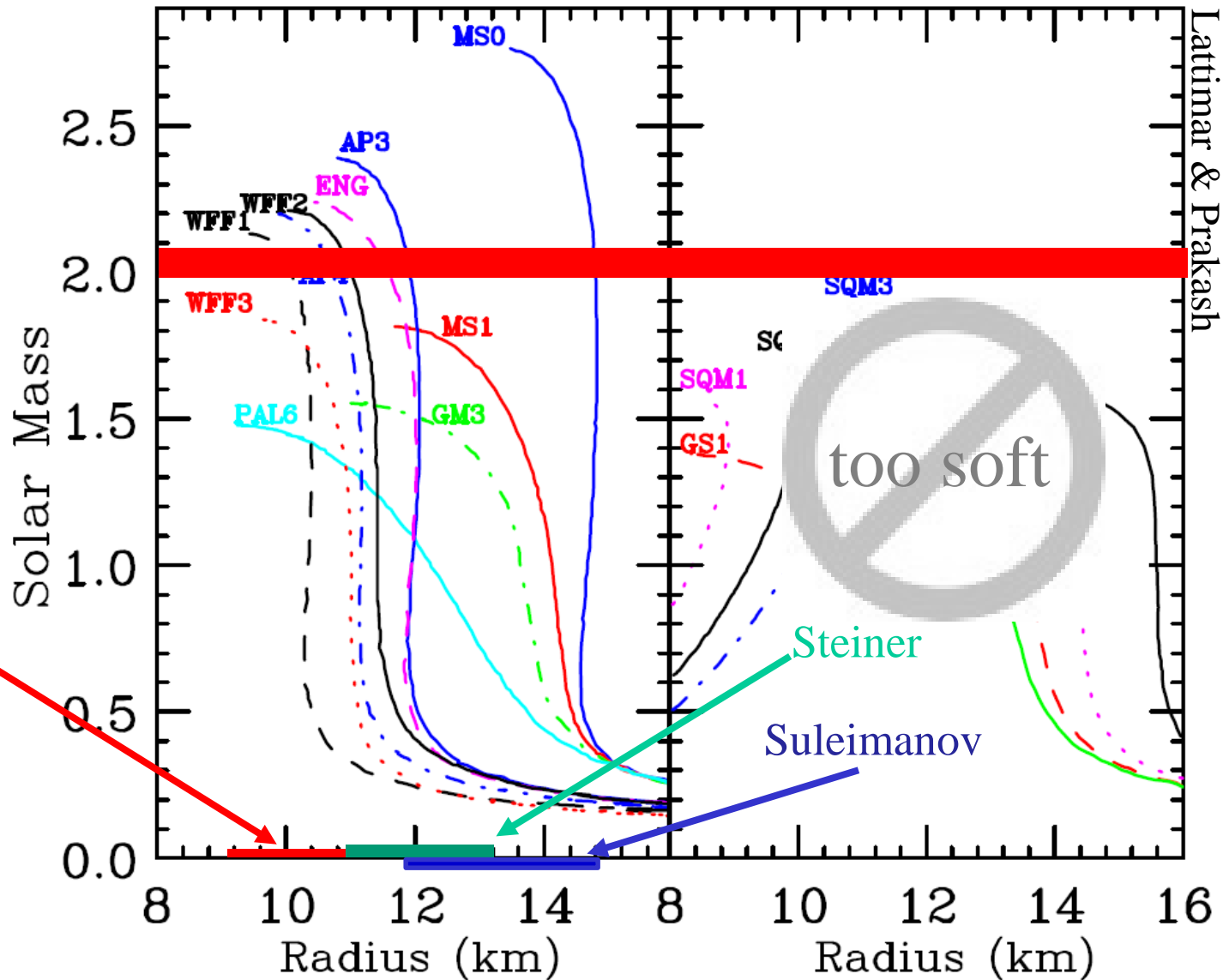
$$\frac{dP}{dr} = -G \frac{\mathcal{E}(r)M(r)}{r^2} \left[1 + \frac{P(r)}{\mathcal{E}(r)} \right]$$

$$\left[1 + \frac{4\pi r^3 P(r)}{M(r)} \right] \left[1 - \frac{2GM(r)}{r} \right]^{-1}$$



Recent observations of Neutron Stars (radius/Radii)

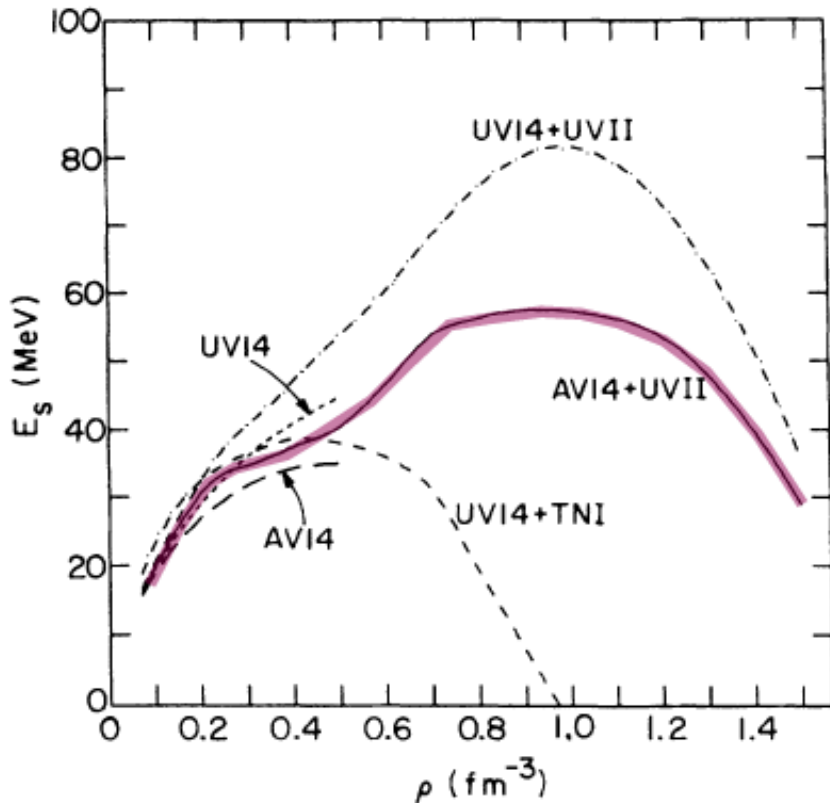
S. Guillot, et al *Astrophys. J.*
772, 7 (2013), 1302.0023



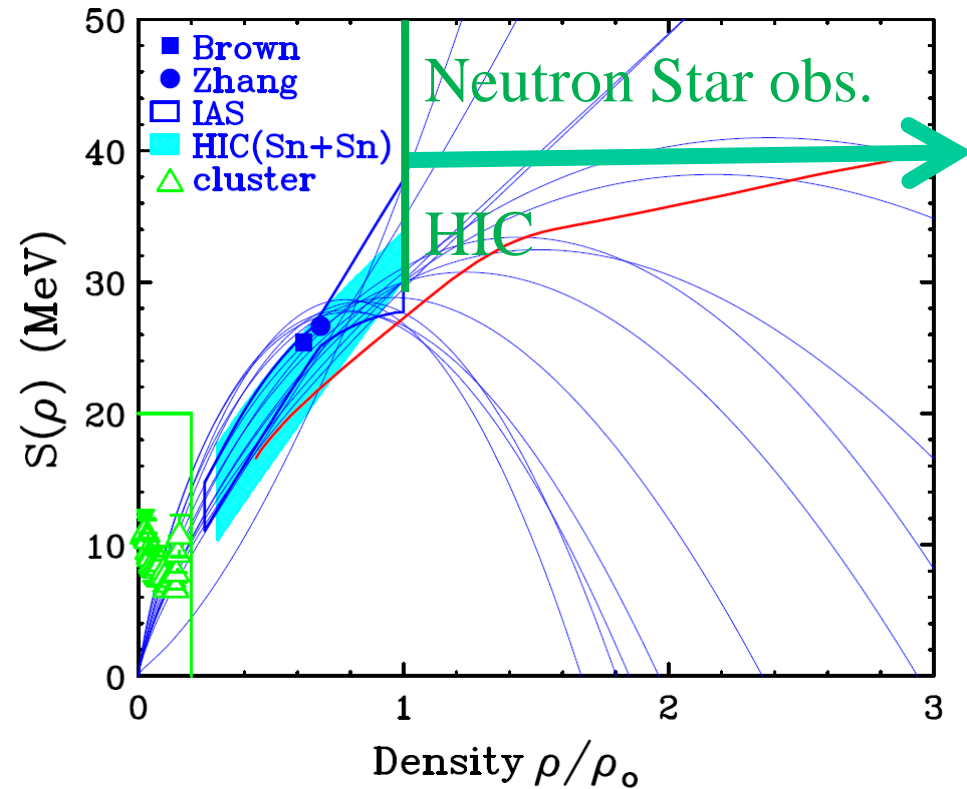
Very small Neutron Star radius rules out nearly all EOS

Density dependence of symmetry energy at supra-saturation density

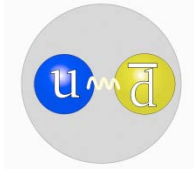
Wiringa, Fiks, & Fabrocini 1988



Skyrme interactions



Above saturation density, the symmetry energy density dependence may have a different energy dependence than Skyrme interactions.



Pion Observable

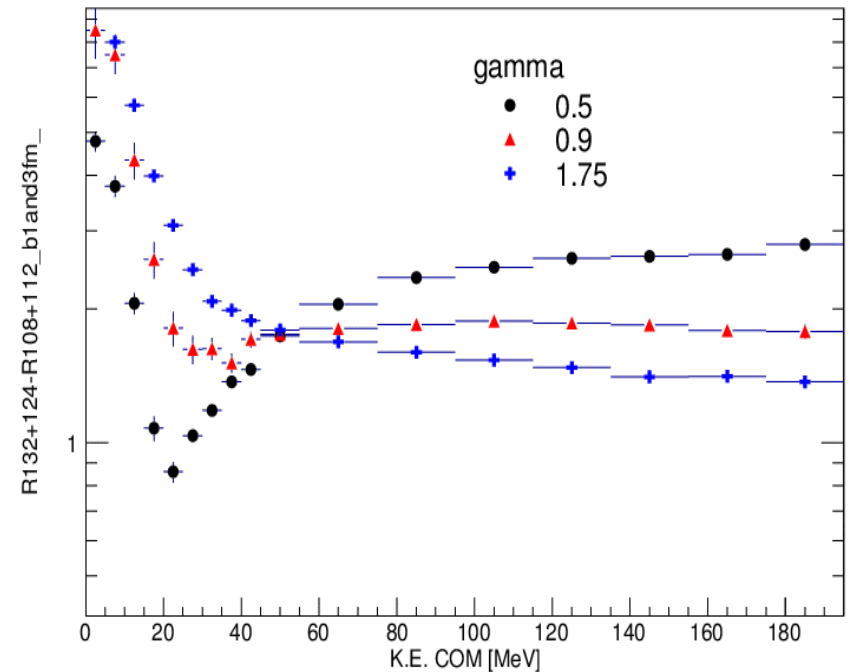
Resonance	Quarks	Δ Formation	Δ Decay
Δ^{++}	uuu	$p + p \rightarrow n + \Delta^{++}$	$\Delta^{++} \rightarrow \pi^+ + p$
Δ^+	uud	$p + p \rightarrow p + \Delta^+$	$\Delta^+ \rightarrow \pi^+ + n$
Δ^0	udd	$n + n \rightarrow n + \Delta^0$	$\Delta^0 \rightarrow \pi^- + p$
Δ^-	ddd	$n + n \rightarrow p + \Delta^-$	$\Delta^- \rightarrow \pi^- + n$

Pros:

- Produced in direct n p collisions – sensitive to symmetry energy
- exit in early time

Cons:

- Cross-section is low and
- Easily reabsorbed in collision medium



- Pion ratios are most sensitive compared to n/p or t/3He ratios
- Differences of pion spectra are more sensitive than ratios of integrated yields.
- A new detector is needed to probe these observables

New Detector

Radioactive Beam:

- low luminosity → large coverage

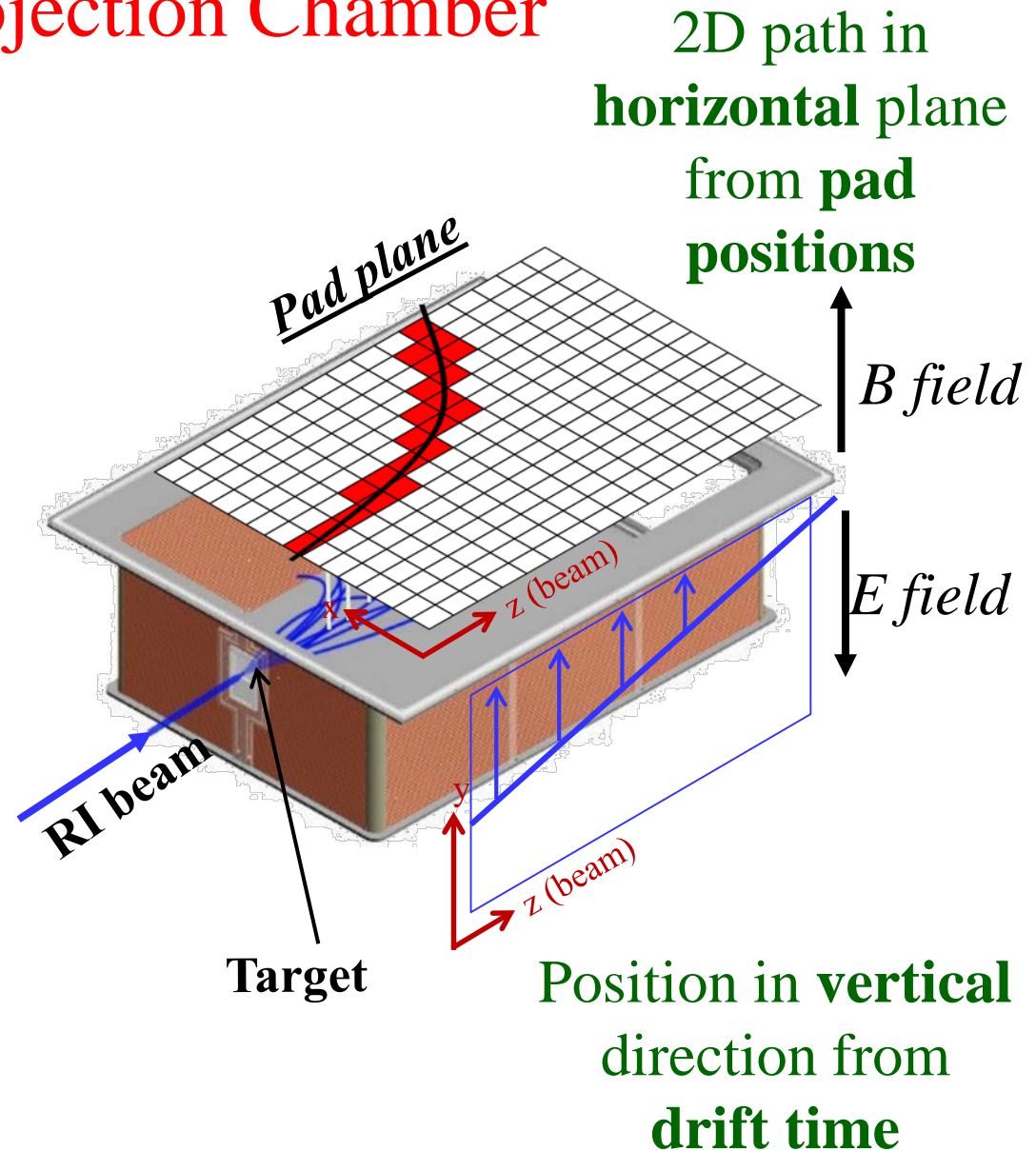
High Resolution:

- resolve many different species of produced particles
- distinguish particles by mass and charge (π^+ , π^-)
- track particles in an applied magnetic field

Versatile for a wide range of experimental programs

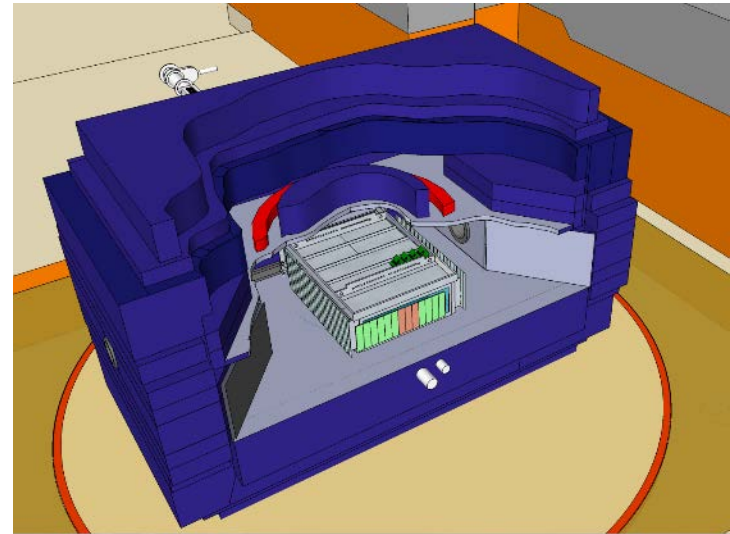
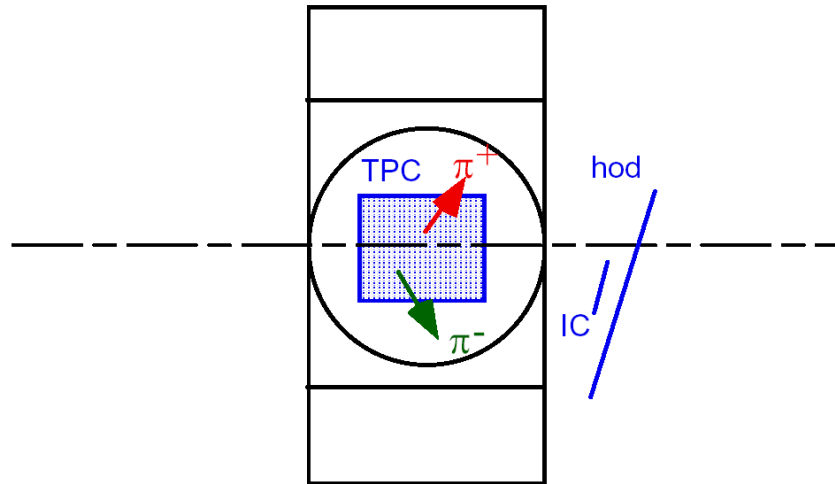
Time-Projection Chamber

- Products from reaction ionize detector gas inside a field cage
- Electron signal is amplified by a wire plane
- The time at which the electrons hit the pads provides the third dimension

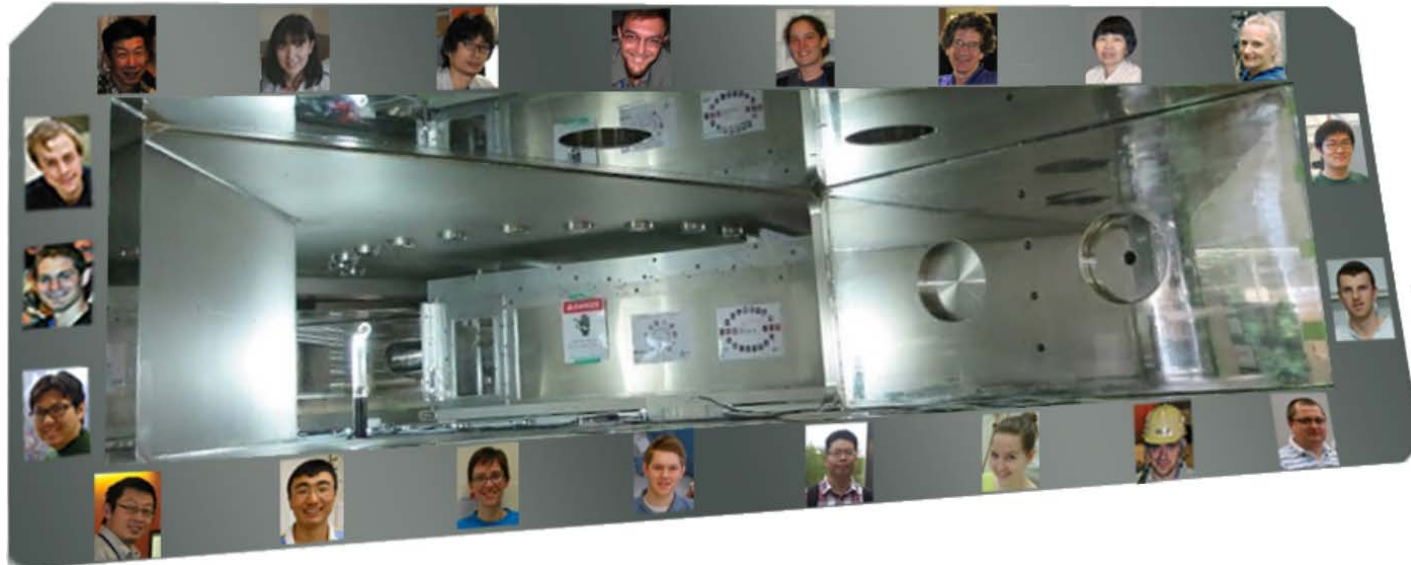


2006 March RIKEN workshop, by Bill Lynch

Possible RIKEN (EOS) program



July 23, 2014



Joint US-Japan project

US Collaboration:

December/2008: submit DOE FOA proposal for \$1.2 M

November/2009: Proposal approved (CAGRA, SAMURAI-Si, JUSEIPEN)

October/2010: Project start date; Construction & Shipment of TPC, and travel (help from JUSEIPEN)

Japanese collaboration: NEXT

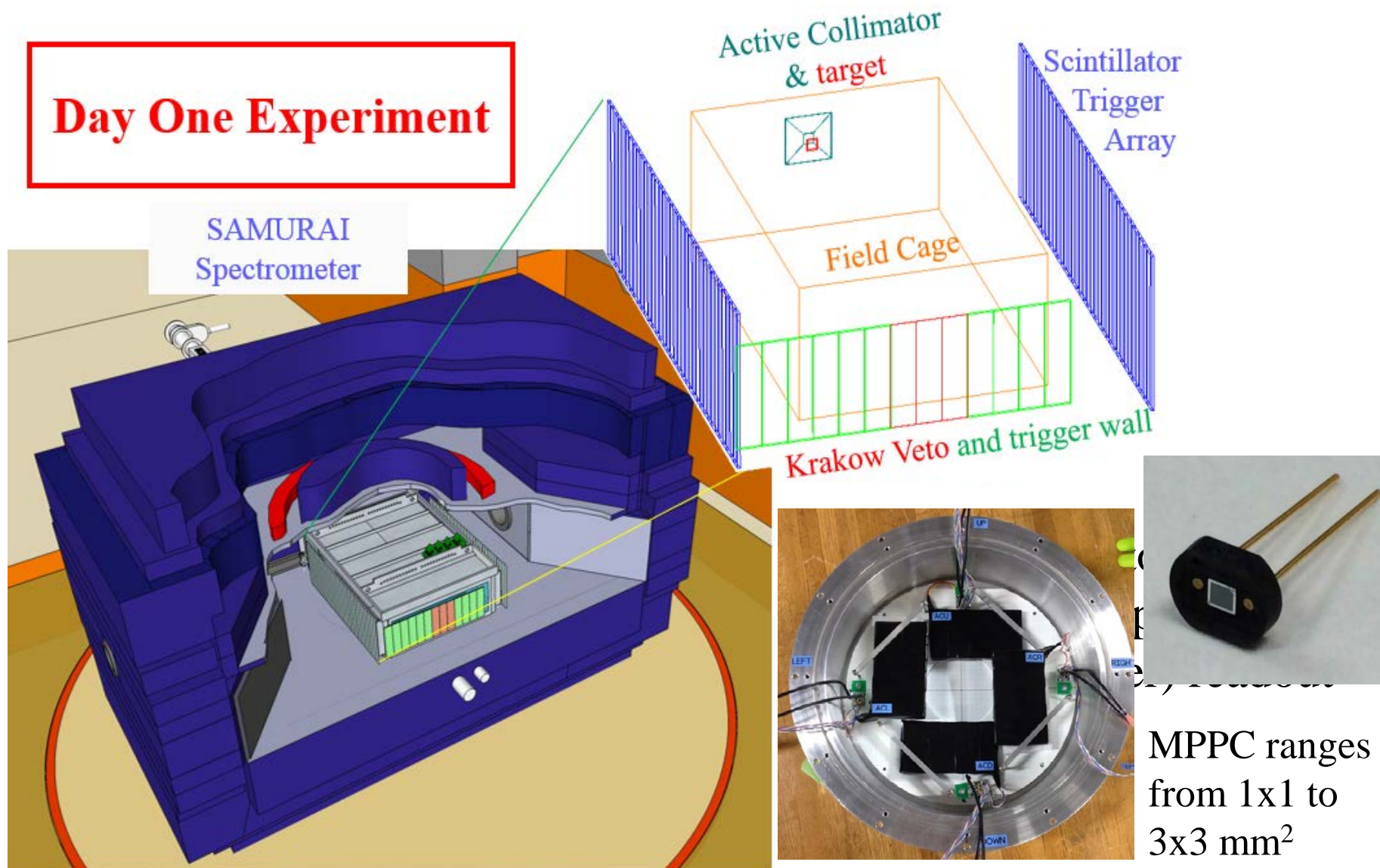
Part of “Material Science of Quarks”

~100 M yen for TPC electronics, Ancillary trigger scintillation array, Targets, TPC gas handling system, TPC laser calibration system, Data acquisition

Approved experiments at RIKEN

Primary	Secondary	Target	δ_{CN}
^{238}U	^{132}Sn	^{124}Sn	0.22
	^{124}Sn	^{112}Sn	0.15
^{124}Xe	^{108}Sn	^{112}Sn	0.09
	^{112}Sn	^{124}Sn	0.15

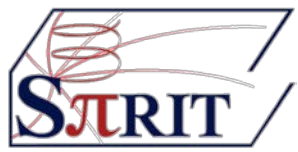
Day 1 experiment: Triggered by multiplicity and beam veto



Yan Zhang, THU

MPPC ranges from 1x1 to 3x3 mm²

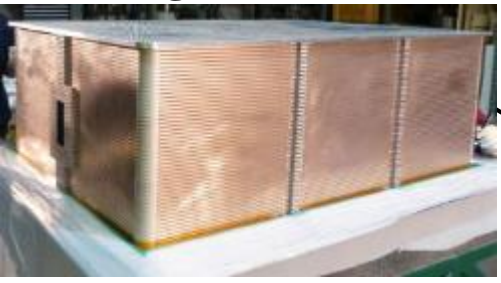
Anatomy of



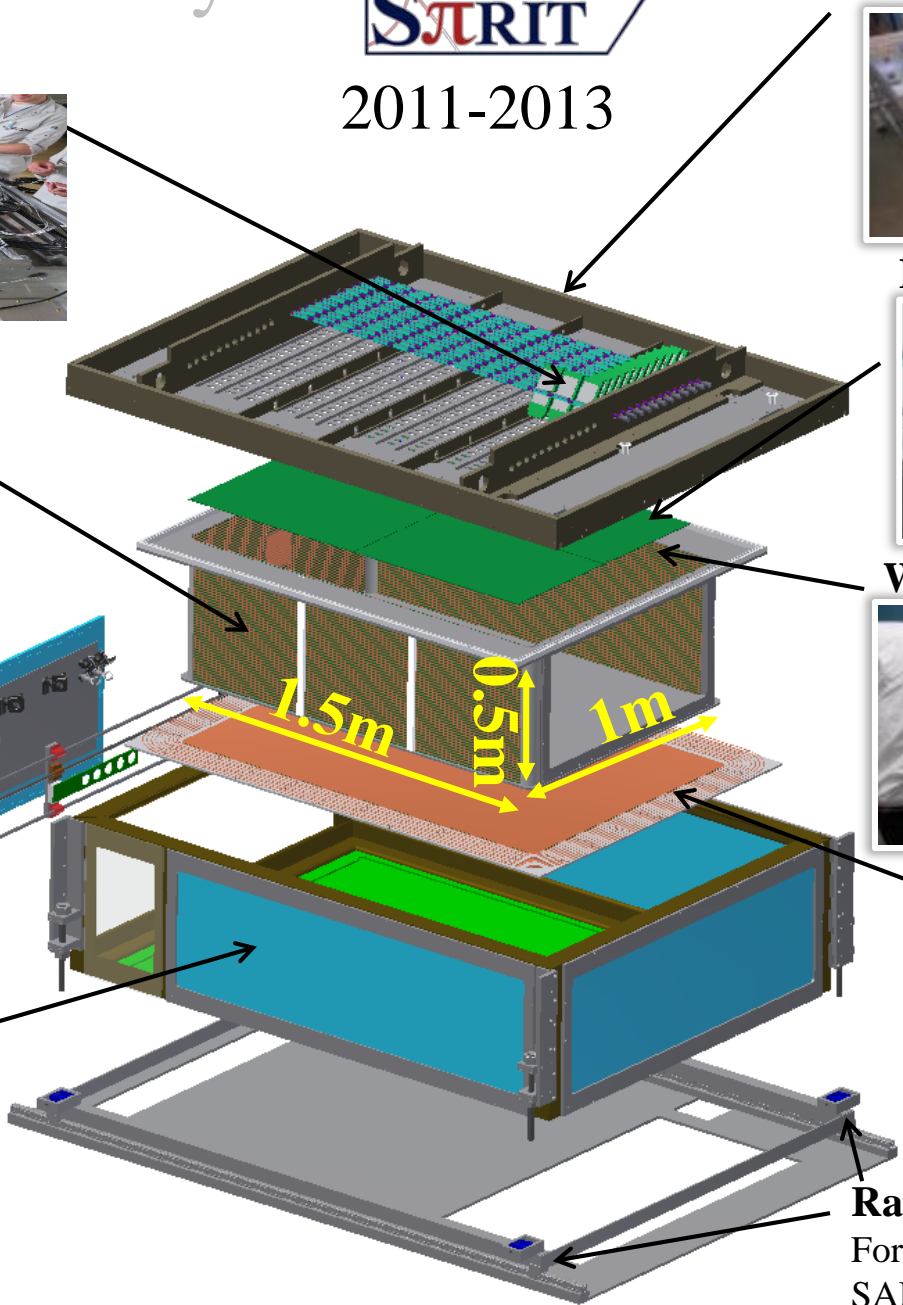
2011-2013



Field Cage



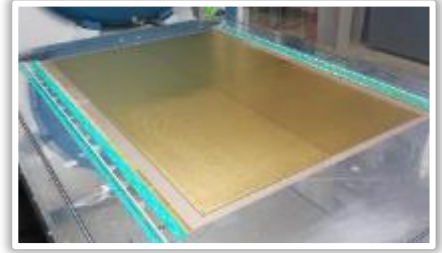
Thin-Walled Enclosure



Rigid Top Plate



Pad Plane (12096 pads)



Wire Planes (e- mult)



Voltage Step-Down



Rails

For inserting TPC into SAMURAI vacuum chamber

Beam

Calibration Laser Optics

Target Mechanism

May, 2013



Feb, 2014



July, 2014



August, 2015

Commissioning of outside SAMURAI

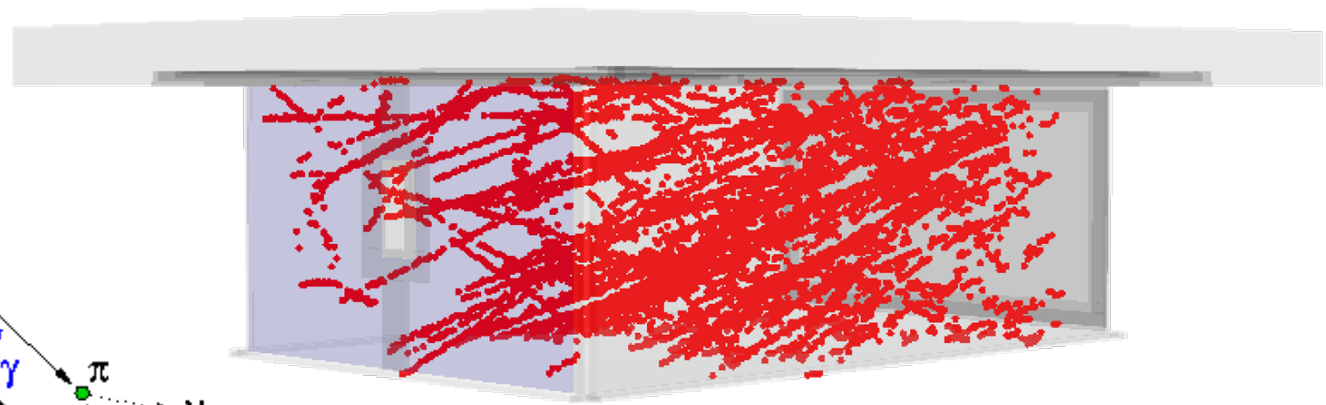
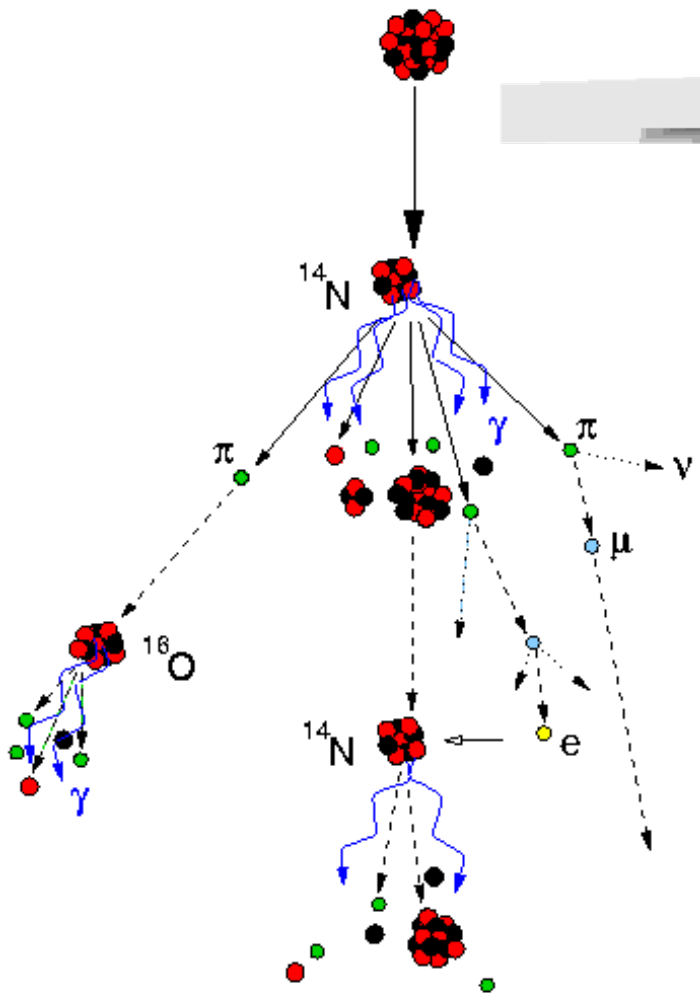
On Site Experimenters (10/23 & 29)



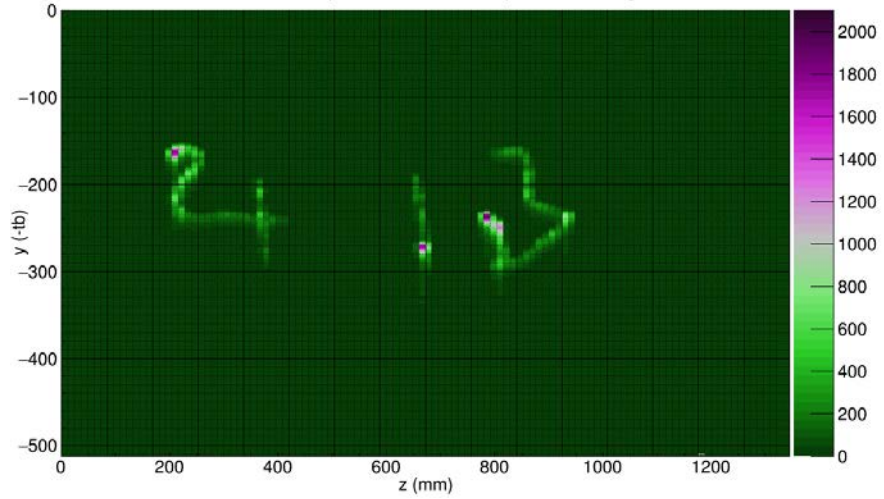
J. Barney (MSU)
G. Cerizza (MSU)
J. Estee (MSU)
*T. Isobe (RIKEN)**
G. Jhang (Korea U)
M. Kaneko (Kyoto U)
Y. Kim (RISP)
M. Kurata-Nishimura (RIKEN)
P. Lasko (IFN, Krakow)
H. Lee (RISP)
J. Lee (Korea U)
J. Lukasik (IFN, Krakow)
*W. Lynch (MSU)**
*T. Murakami (Kyoto U)**
P. Pawlowski (IFN, Krakow)
K. Pelczar (IFN, Krakow)
C. Santamaria (MSU)
D. Suzuki (RIKEN)
*B. Tsang (MSU)**
Y. Zhang (Tsinghua U)

**spokesperson*

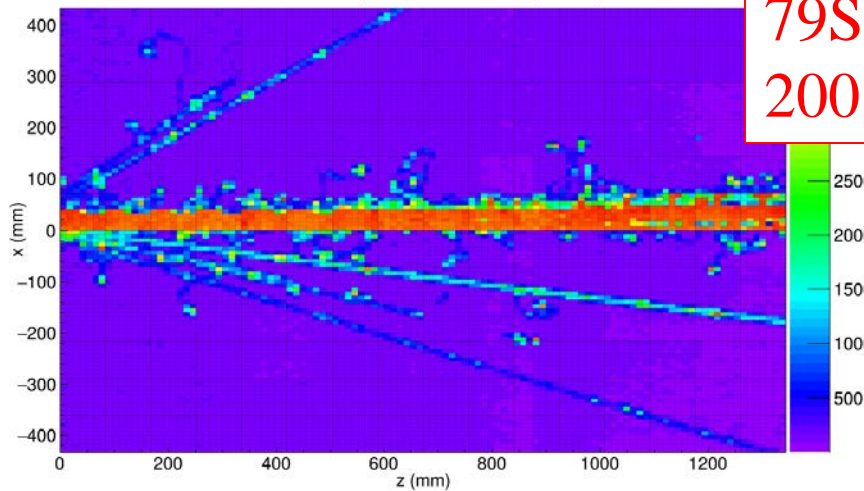
Commissioning of Cosmic Events October, 2015



Event ID: 24 (Gain calibrated) - Beam right view

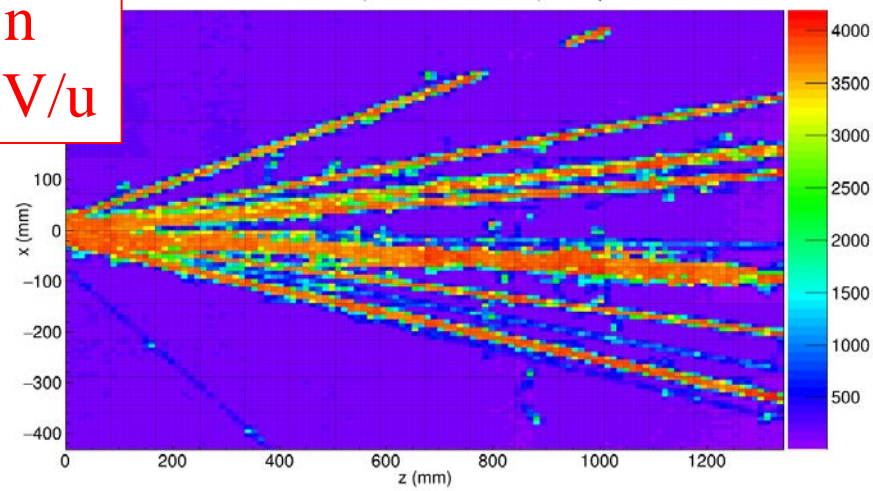


Event ID: 39 (Gain calibrated) - Top view



**$^{79}\text{Se}+\text{Sn}$
200 MeV/u**

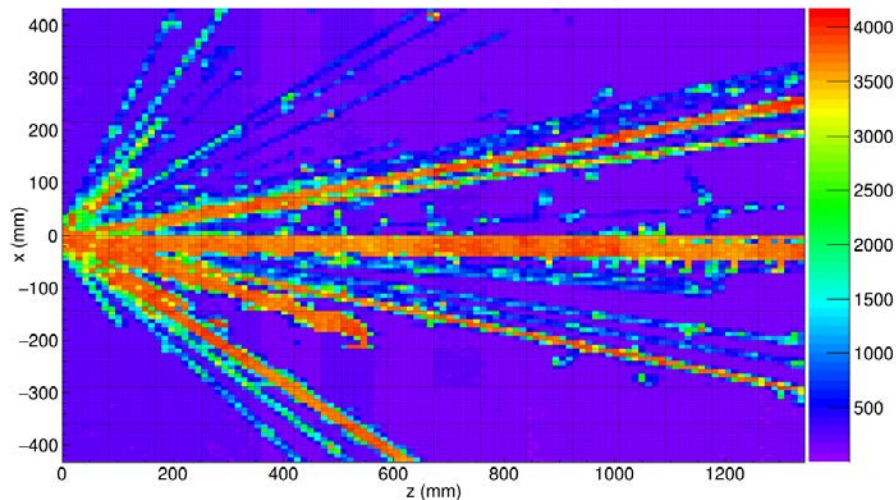
Event ID: 82 (Gain calibrated) - Top view



Event from Kyoto multiplicity > 0
trigger + beam veto

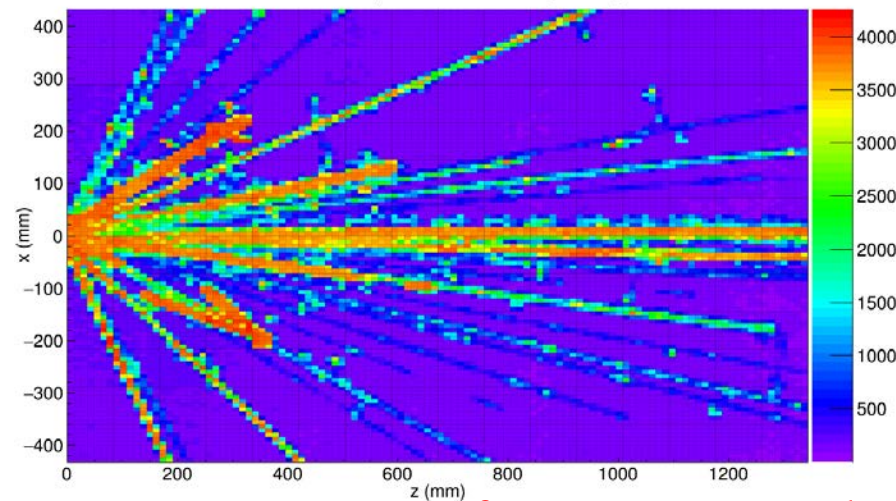
multiplicity > 1

Event ID: 87 (Gain calibrated) - Top view



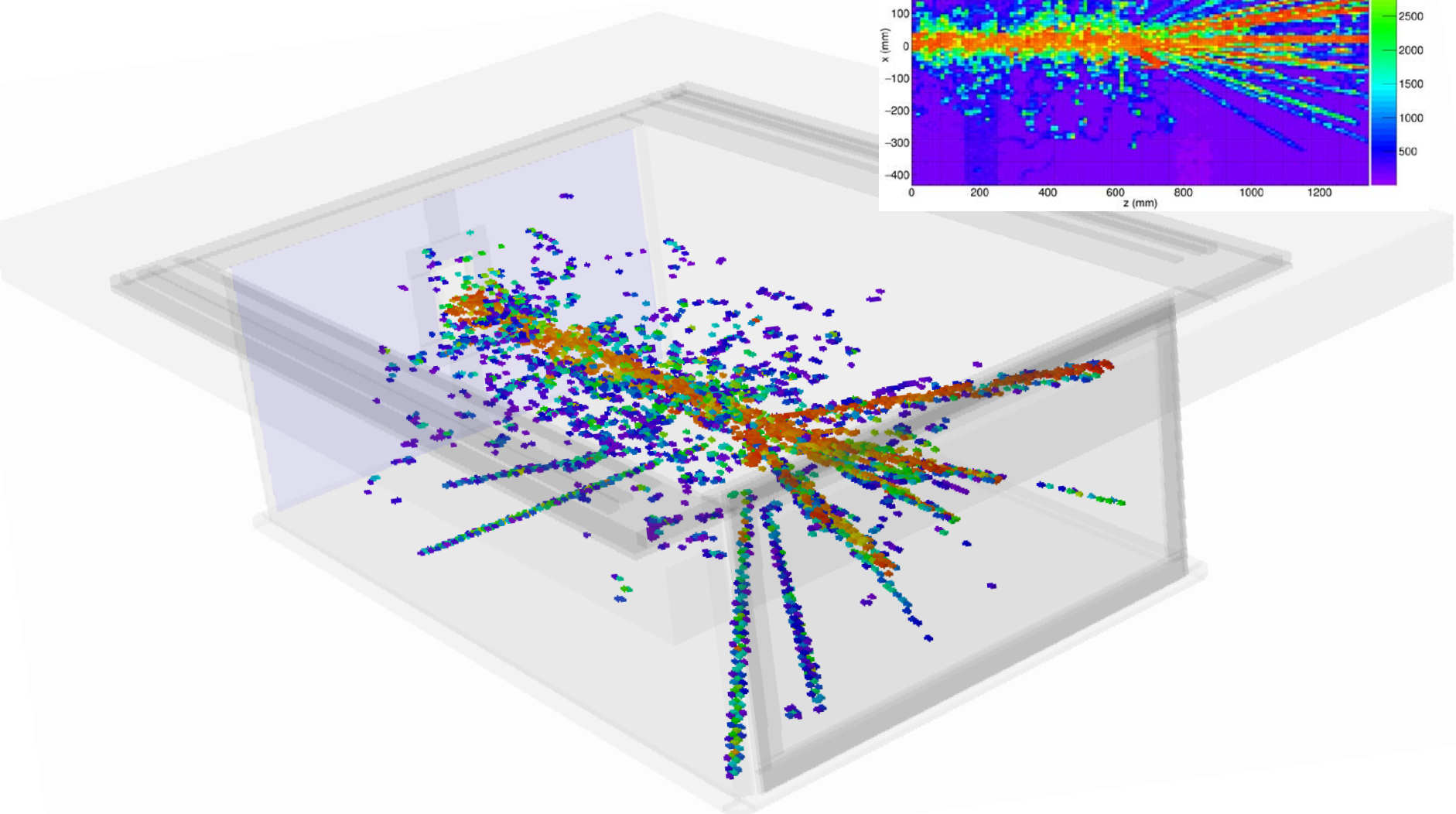
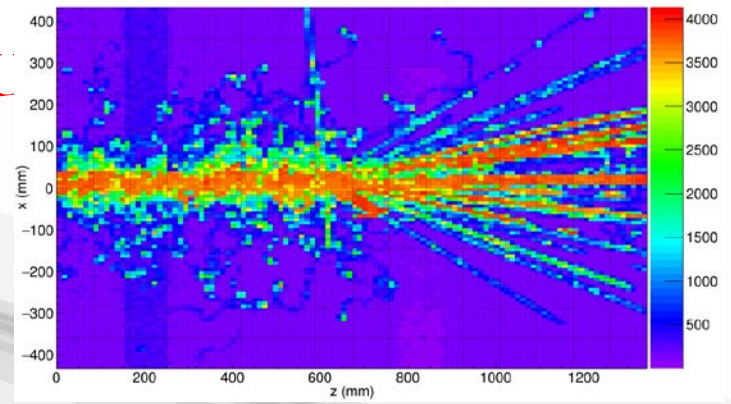
multiplicity > 2

Event ID: 17 (Gain calibrated) - Top view



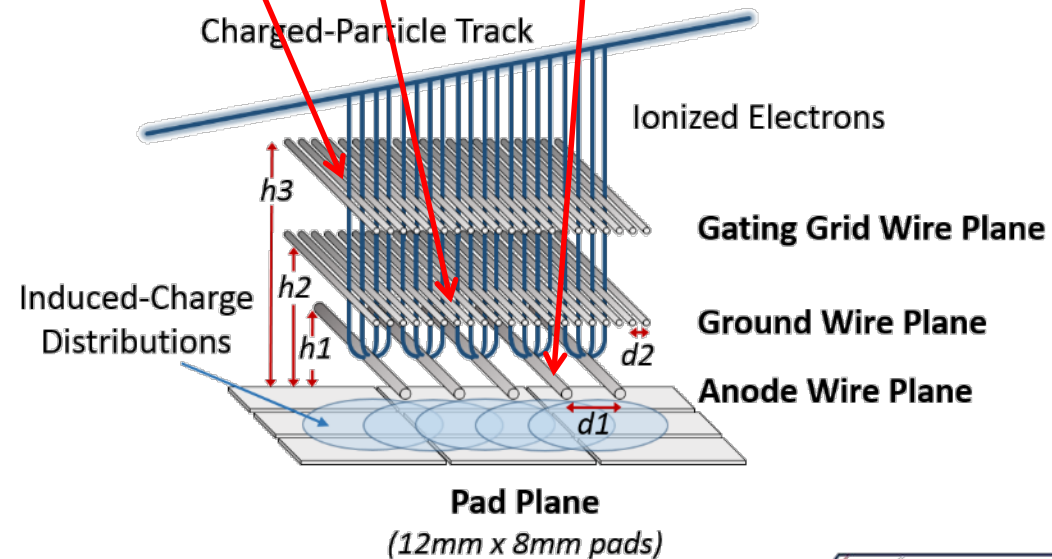
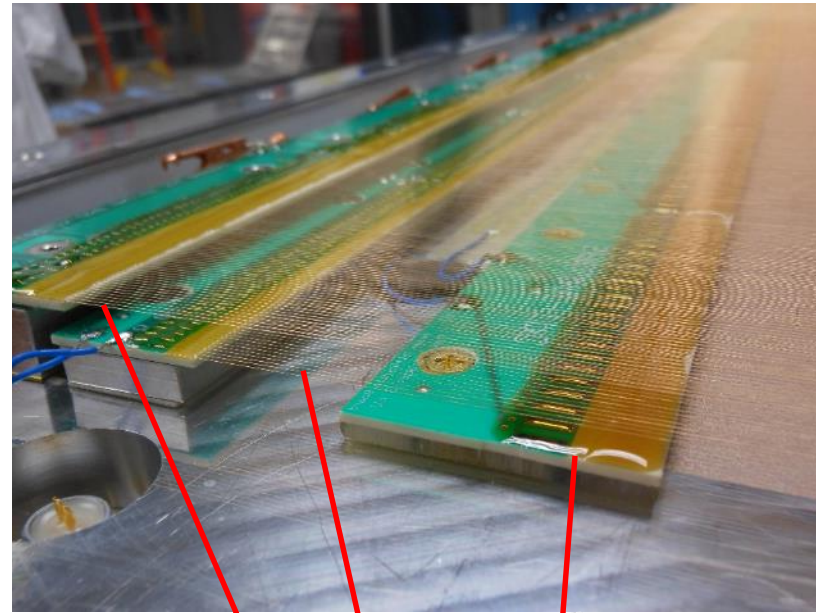
Event from Katana central
trigger + beam veto

Commissioning of outside SAMU



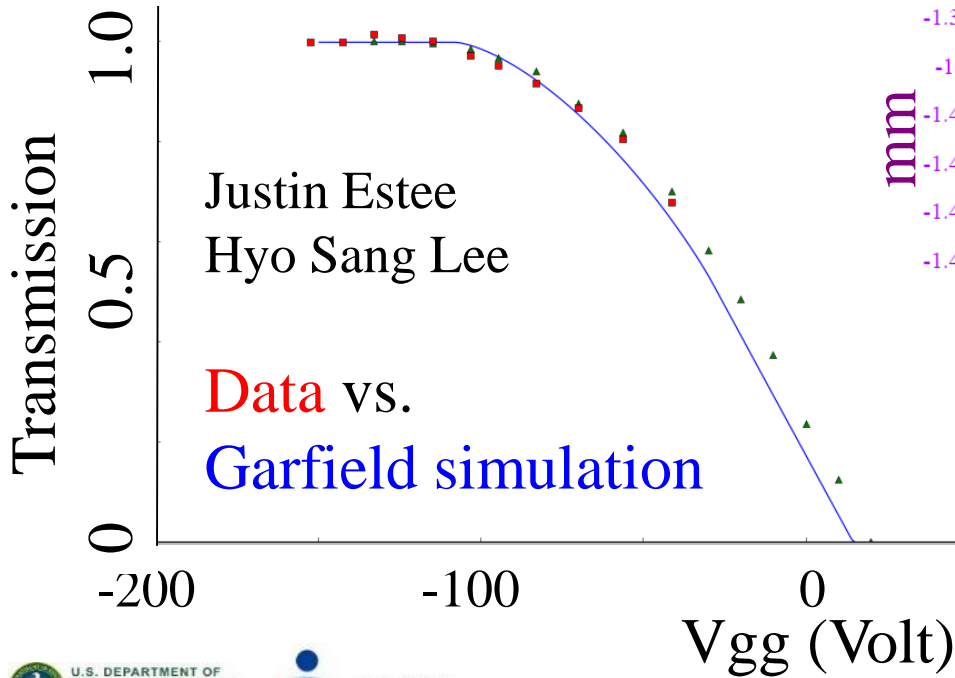
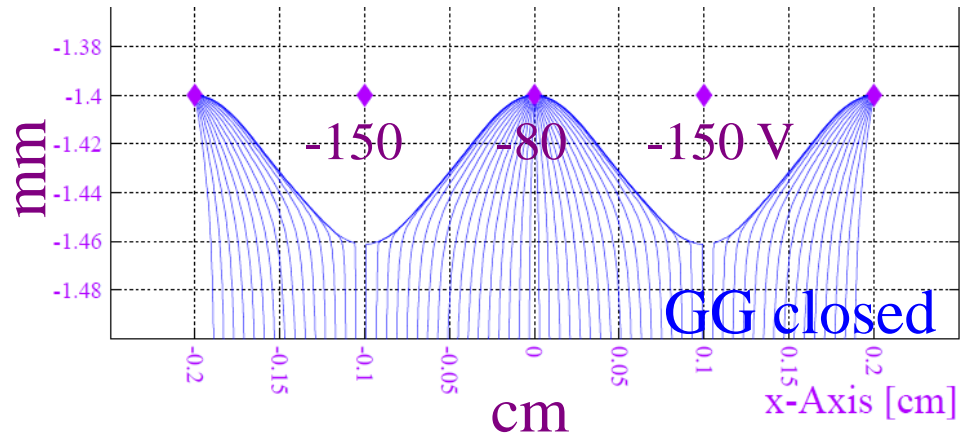
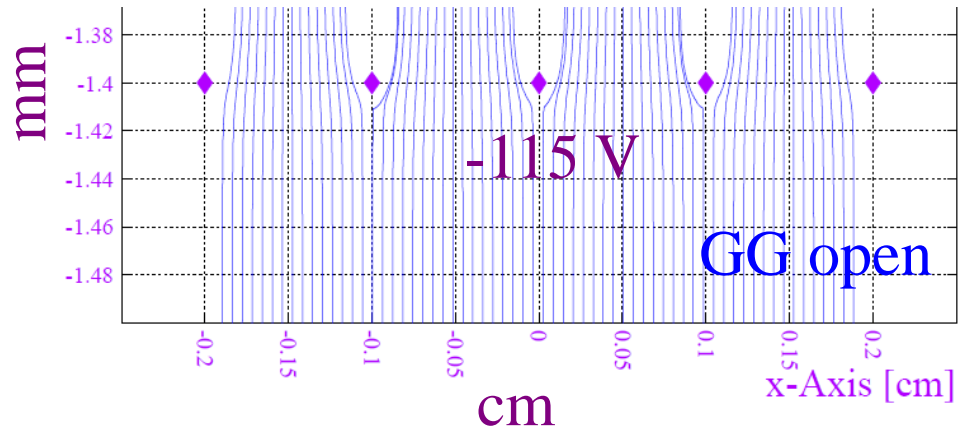
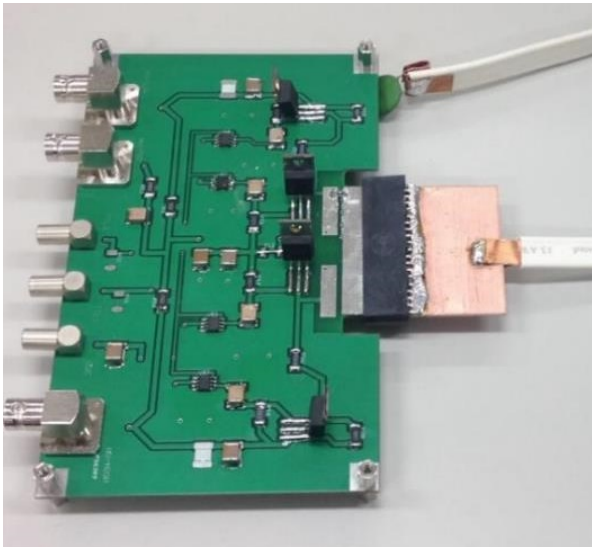
Wire planes

- Anode and ground plane creates avalanche region for electrons
- Anode plane induces image charge on the pad plane
- Gating grid closes off amplification region when not triggered



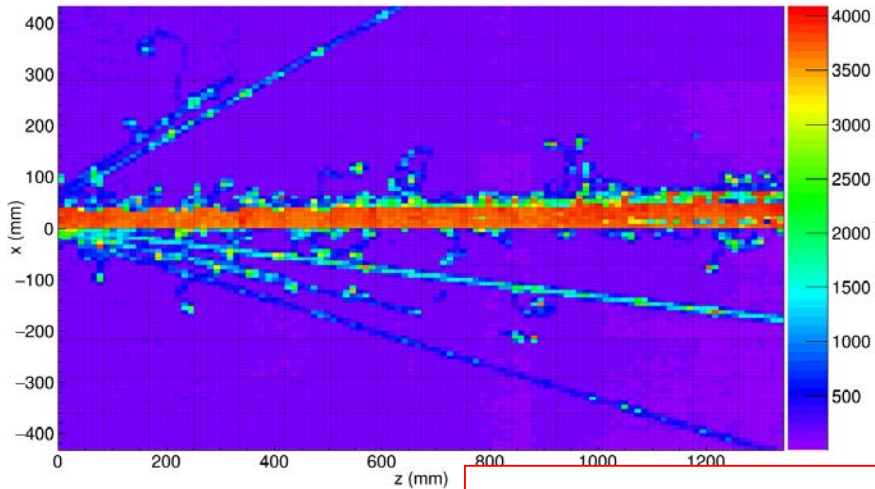
Plane	height (mm)	pitch (mm)	diameter(μm)
Anode	4.05	4	20
Ground	8.1	1	75
Gating grid	14	1	75

Gating grid (Suwat Tangwancharoen)

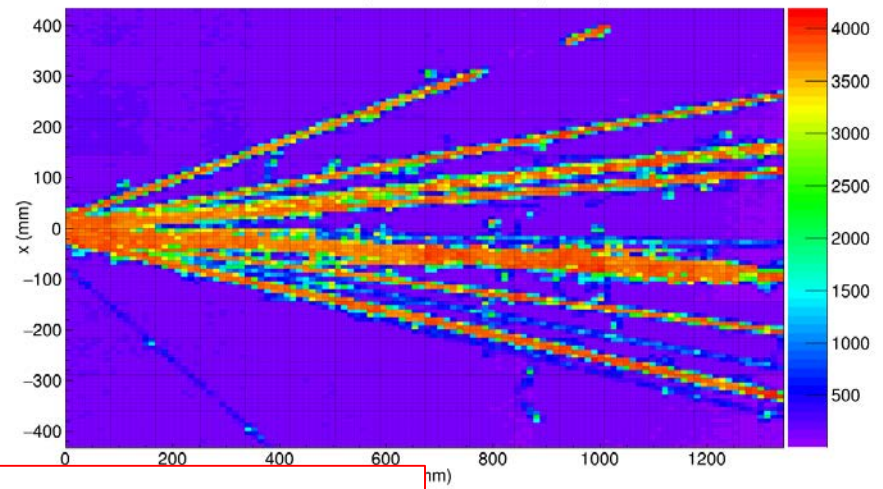


Yao-Feng Zhang, BNU
S. Tangwanchoren
J. Justin

Event ID: 39 (Gain calibrated) - Top view



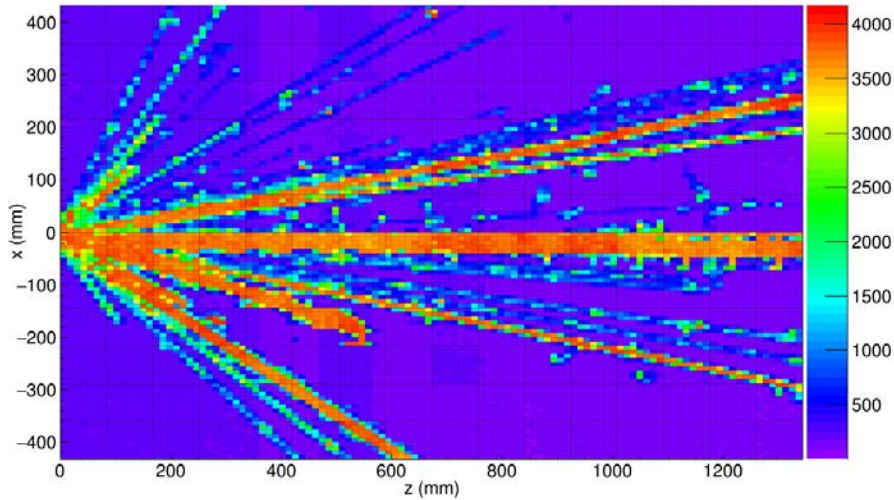
Event ID: 82 (Gain calibrated) - Top view



Event from Kyoto
trigger + beam veto

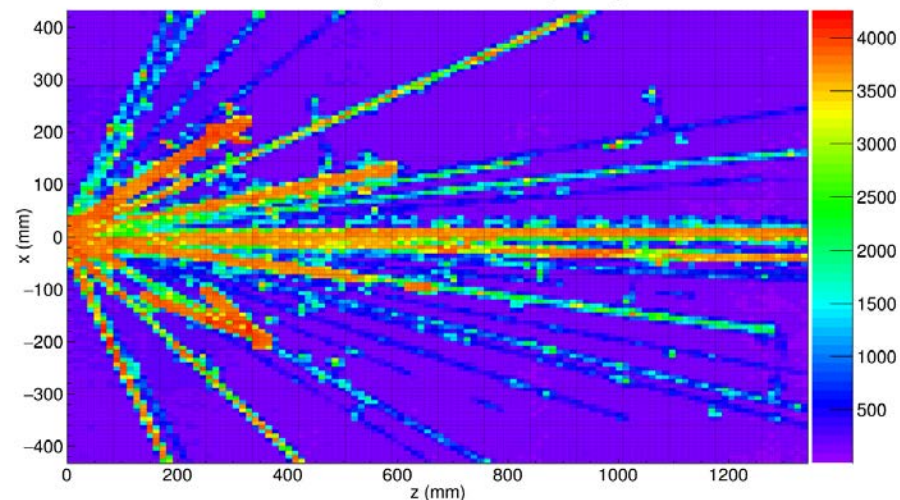
Challenge I: Analysis of Commissioning Data (4 TB)

Event ID: 87 (Gain calibrated) - Top view



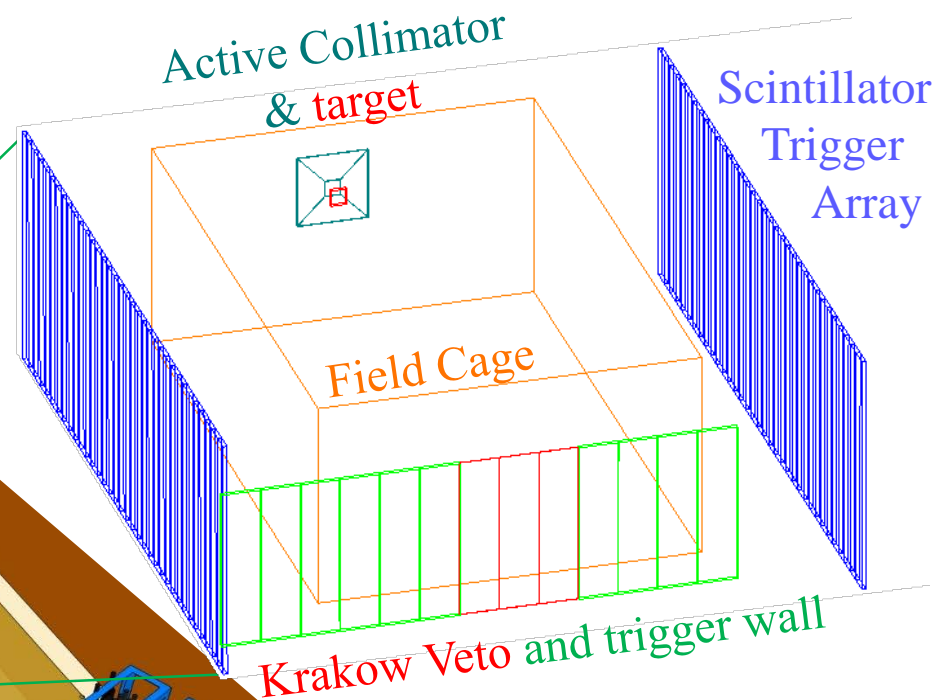
multiplicity > 2

Event ID: 88 (Gain calibrated) - Top view

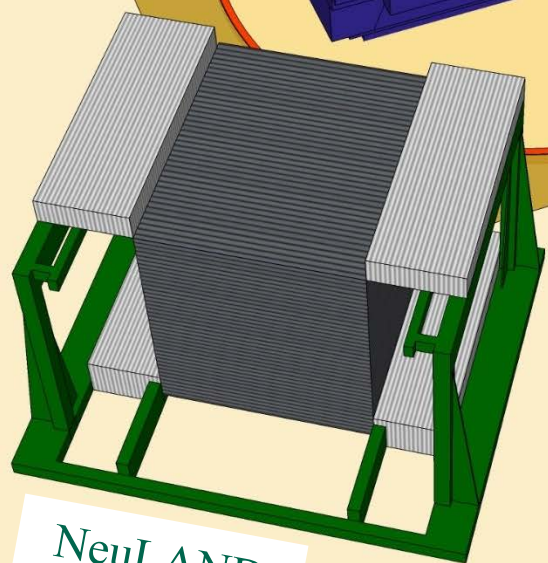
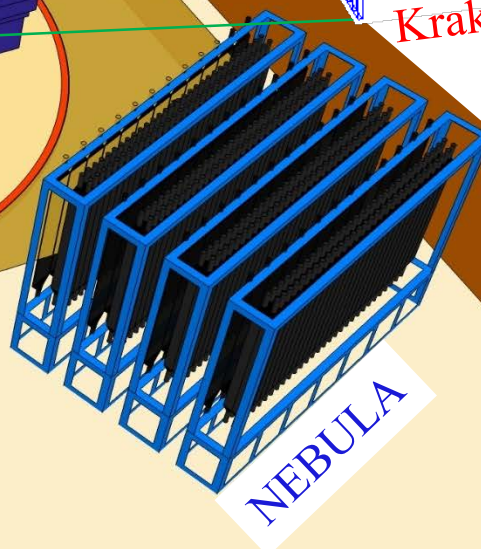
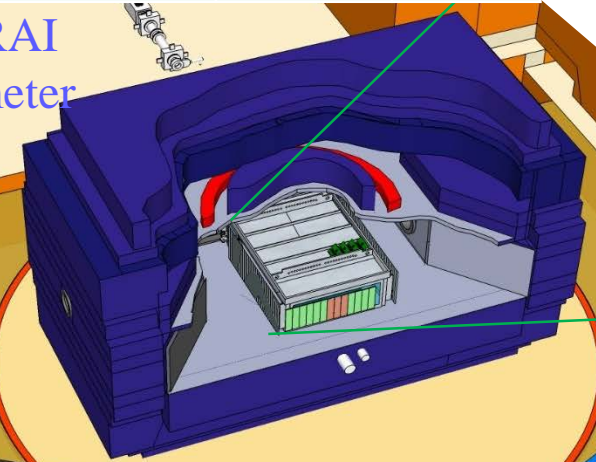


Event from Katana central
trigger + beam veto

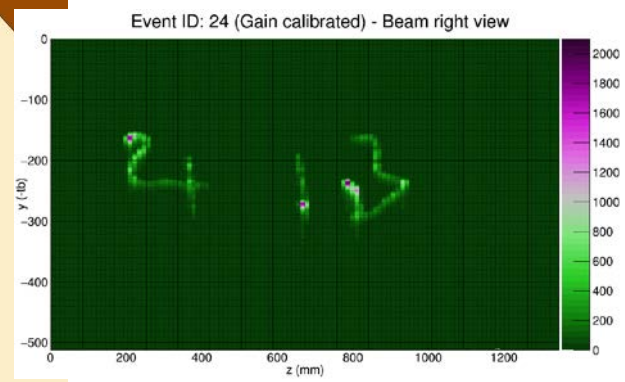
Primary	Secondary	Target	δ
^{238}U	^{132}Sn	^{124}Sn	0.22
	^{124}Sn	^{112}Sn	0.15
^{124}Xe	^{108}Sn	^{112}Sn	0.09
	^{112}Sn	^{124}Sn	0.20

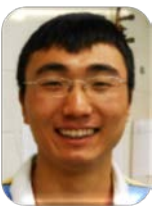


SAMURAI Spectrometer



Challenge II: Completion of proposed experiments





U.S. DEPARTMENT OF
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MEXT
MINISTRY OF EDUCATION,
CULTURE, SPORTS,
SCIENCE AND TECHNOLOGY-JAPAN

Commissioning of outside SAMURAI

On Site Experimenters
(10/23 & 29)



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G. Cerizza (MSU)
J. Estee (MSU)
*T. Isobe (RIKEN)**
G. Jhang (Korea U)
M. Kaneko (Kyoto U)
Y. Kim (RISP)
M. Kurata-Nishimura (RIKEN)
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Y. Zhang (Tsinghua U)

**spokesperson*

Welcome to

6th international symposium on nuclear symmetry energy (NuSym2016)

Tsinghua University, Beijing, Jun. 13-17, 2016



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