

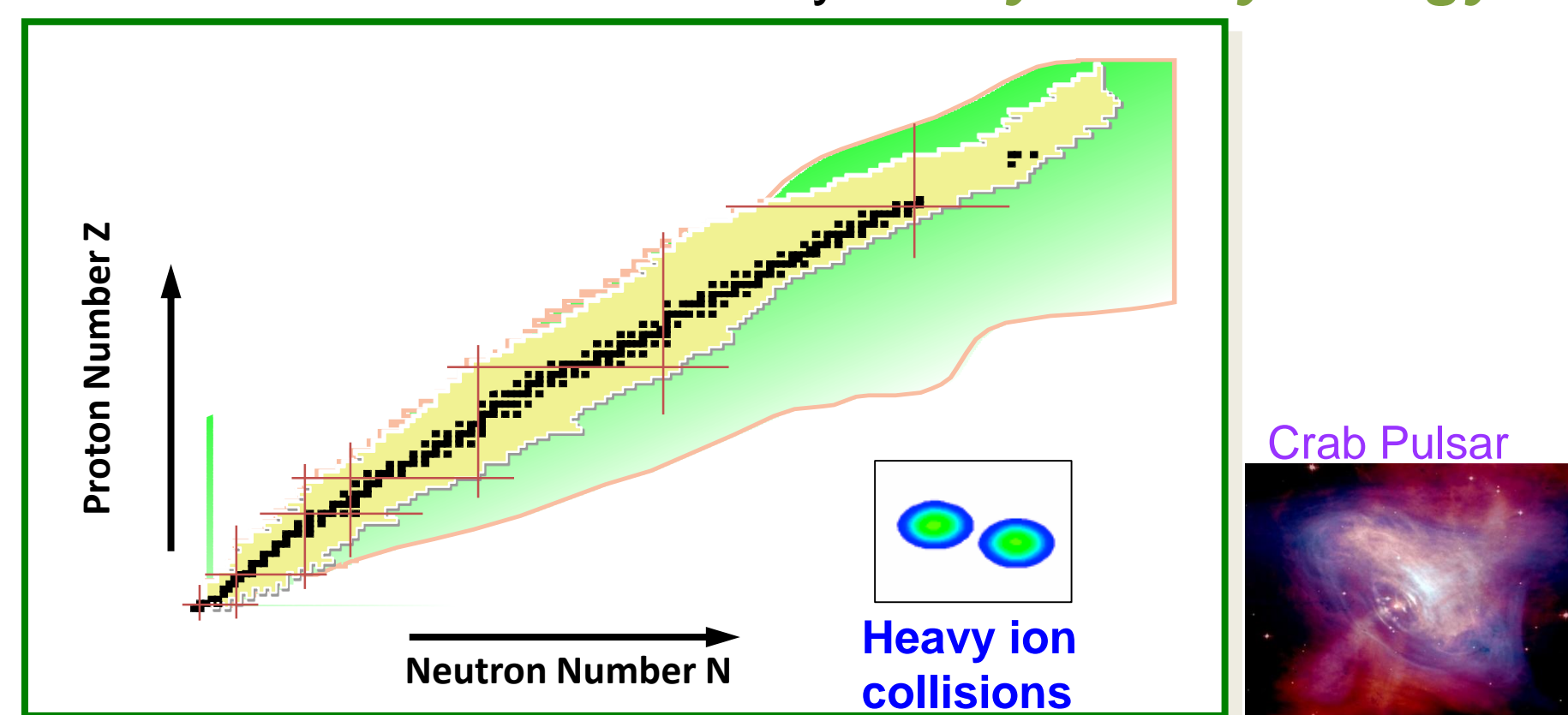
HiRA Array Upgrade to Detect Higher Energy Fragments from Heavy Ion Collisions

Corinne Anderson, Michigan State University

Advisors: William Lynch PhD & Betty Tsang PhD

PHYSICS MOTIVATION

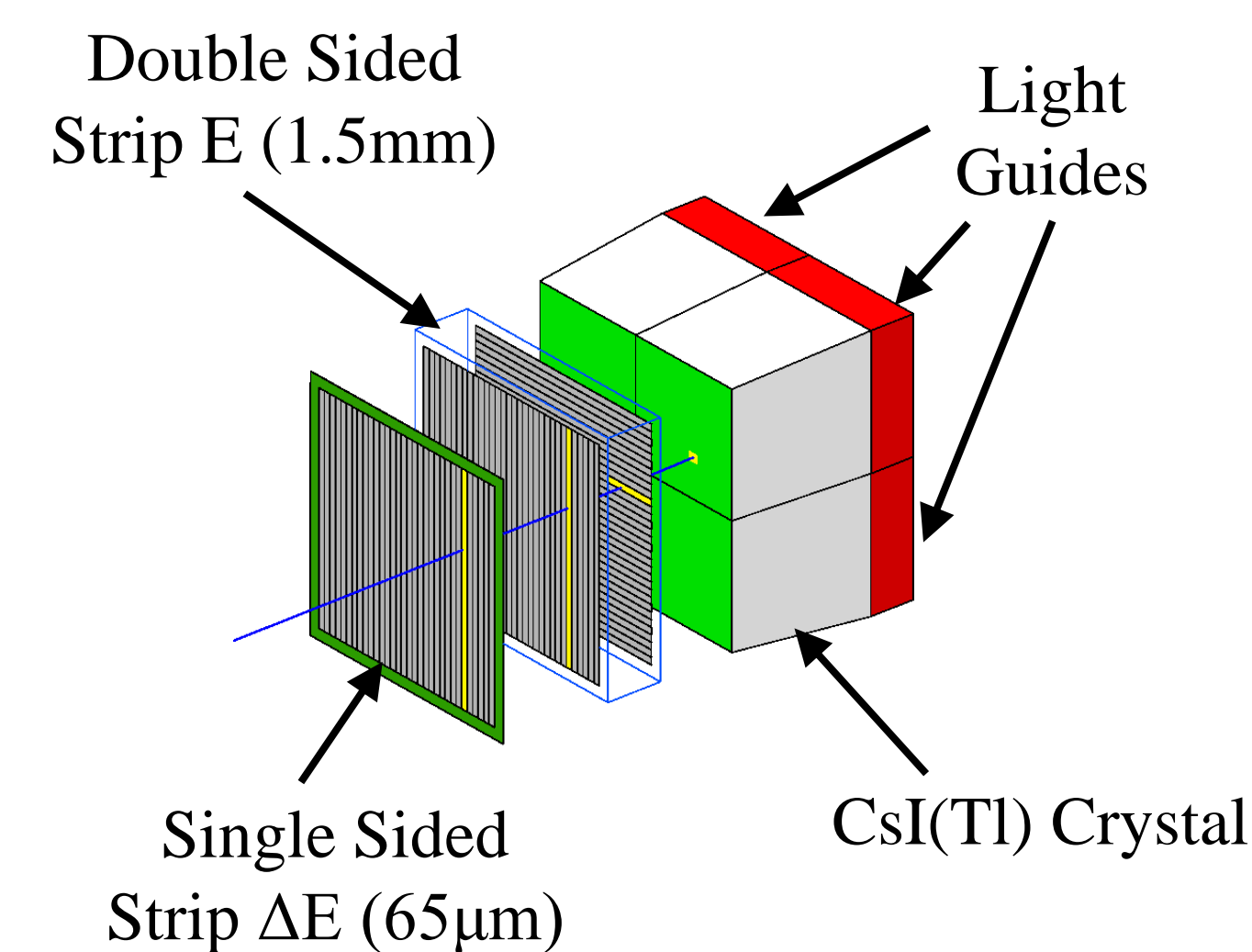
- *Why doesn't neutron star collapse under its own weight (one teaspoon = 5.5 E12 kg)?* Pressure from symmetry energy from the imbalance of neutrons to protons counters the gravitational force.
- The same physics that governs the properties of neutron stars also governs the *properties of nuclei*
- At NSCL, the HiRA group measures fragments, n,p,d,t,3He, alpha, Li up to O isotopes emitted in Sn+Sn collisions to study the *symmetry energy*.



ORIGINAL DEVELOPMENT MOTIVATION

HiRA was an array of **20 charged particle telescopes** designed to provide the **high angular coverage** and **high granularity** required in a variety of experiments.

- Nuclear Structure and Reactions
- Probing the Nuclear EOS & Nuclear Astrophysics
- ~20 papers published



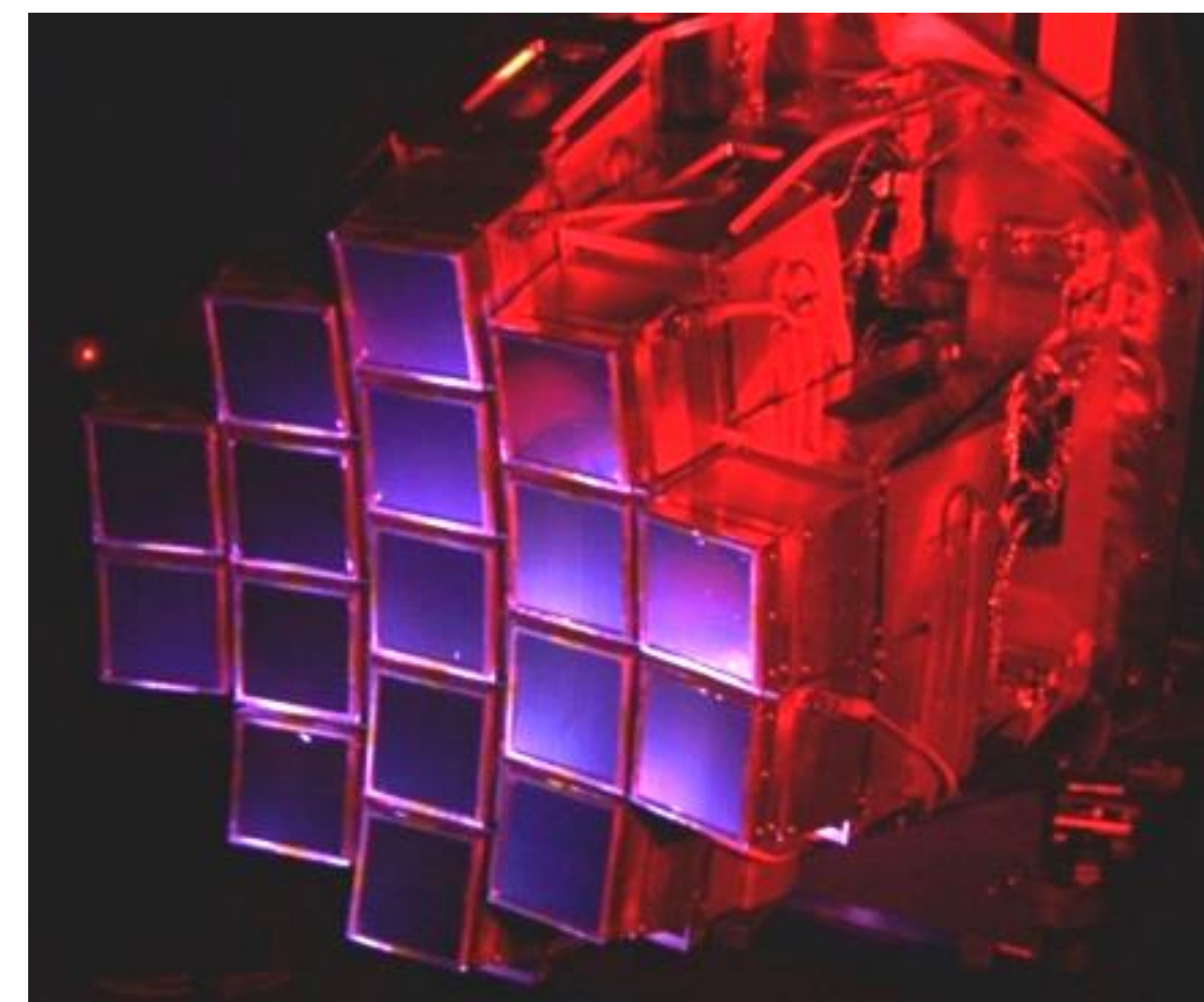
PREVIOUS HIRA TECHNICAL SPECIFICATIONS

- 32 strip, single-sided 65μm "ΔE" layer
- 32 strip, double-sided "E" layer
- Si strip pitch of 2000μm
- **4 CsI(Tl) crystals 4 cm long** for particles that go through Si

PRINCIPLES OF OPERATION

Particle identification: Isotopes are identified using the ΔE-E technique. By comparing the energy loss in one Si detector layer to the energy deposited in another layer, particles are identified based on their mass A and charge Z.

Position sensitivity: The perpendicular strips on the front and back of the thick "E" detector define a pixel that characterizes the particle's trajectory. Each E detector contains 1024 pixels.

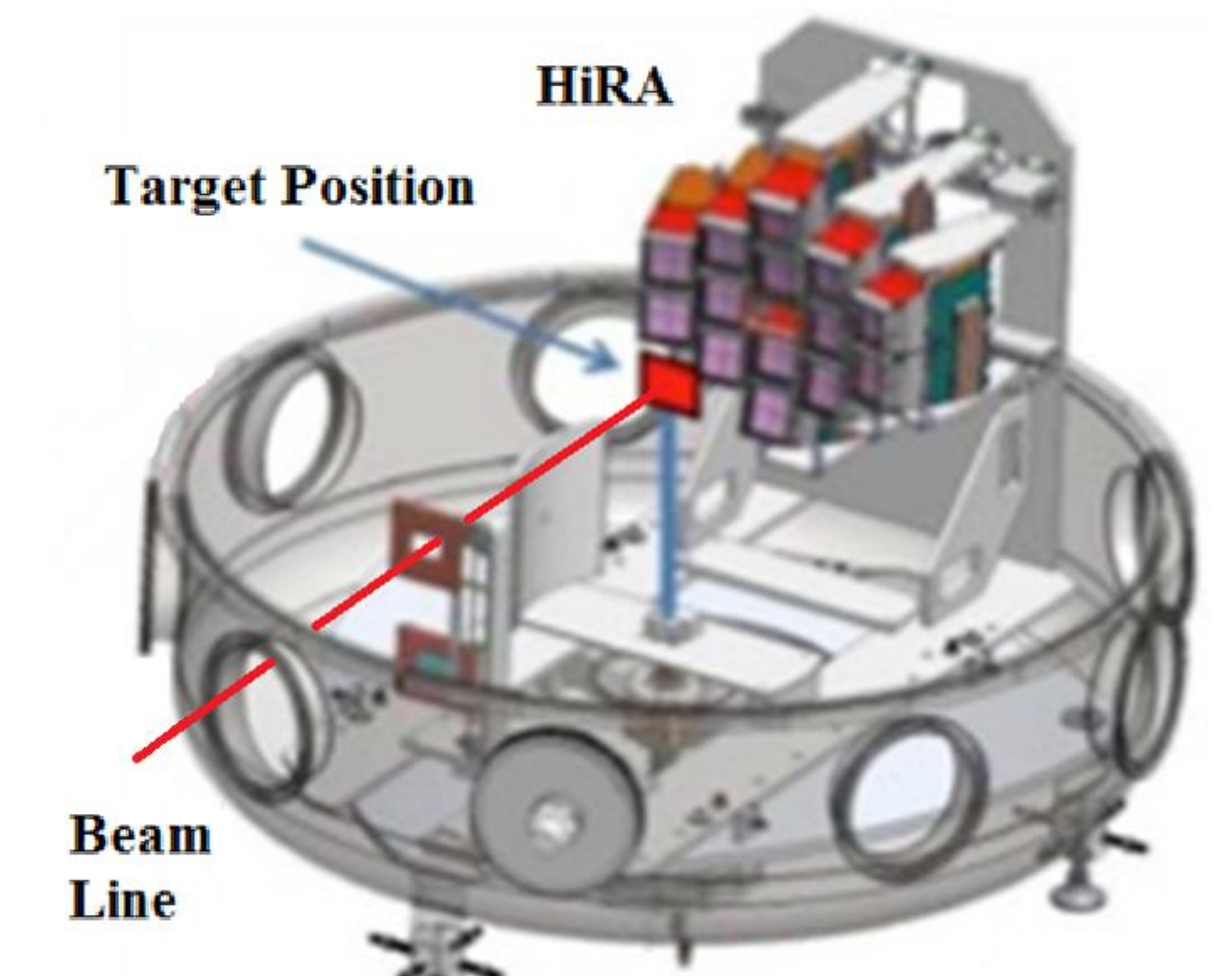


NEW HIRA TECHNICAL SPECIFICATIONS

- Same Si detectors
- **9 CsI(Tl) crystals 10 cm long** for higher energy particles that stop in the CsI crystals after going through Si

CURRENT STATUS (Rough Schedule)

- Conceptual Design: Nearly Complete
- Quotes from Vendor: End of April
- Test single CsI crystal: End of June
- Finalize Crystal Design: Middle of July
- Finalize Preamp Box Design: End of August
- Delivery of upgraded crystal array: End of October
- Assembly: End of December



OUTLOOK

- Finish date: December, 2015
- Approved experiment proposal #14030, expected to run in 2016 & another proposal submitted.

The new telescope design with more crystals of greater length, should identify particles and their positions with **greater sensitivity**. The re-designed telescope array is expected to be used in future experiments.

Motivation for the upgrade

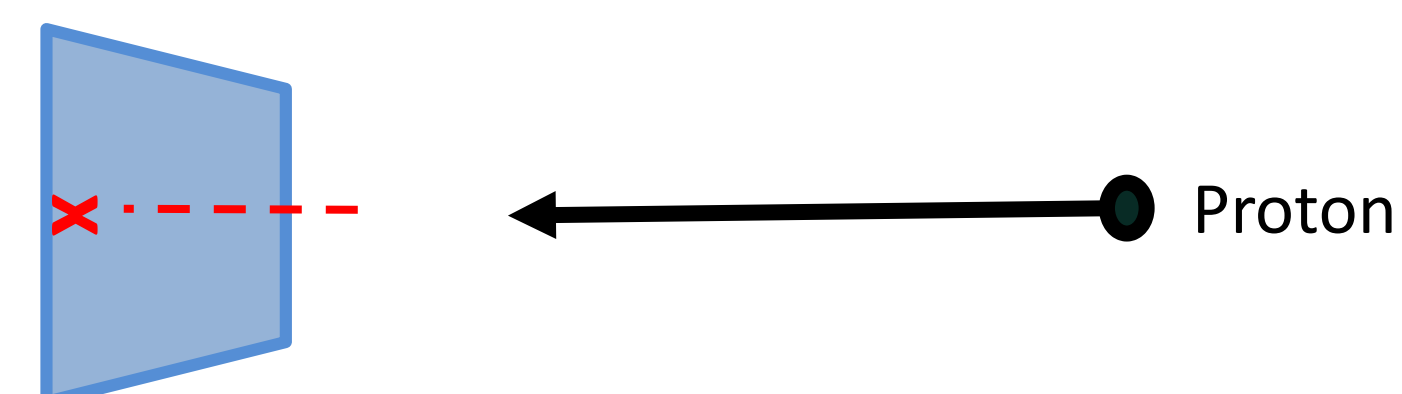
CESIUM IODIDE (CsI) CRYSTALS

CsI crystals are useful for **measuring the energies of nuclear particles** in a reaction. The crystals produce light when electrons are ionized within them as particles travel through. By measuring the light produced, a measurement of the energy of the particles is reached. The Si photo-diodes attached to the crystals have a current \propto to the collected light and to the energy deposited.

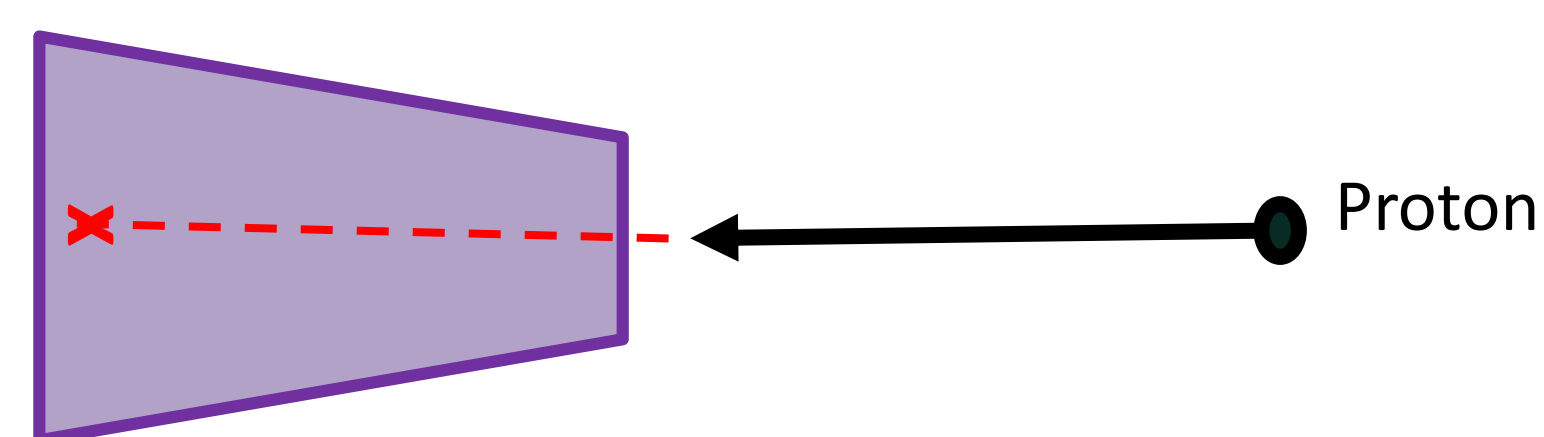
UPDATED DEVELOPMENT MOTIVATION

Detection of **higher energy particles** than previously feasible. 4 cm CsI detect up to 116 MeV proton. HiRA upgrade CsI (10 cm) will detect p up to 190 MeV.

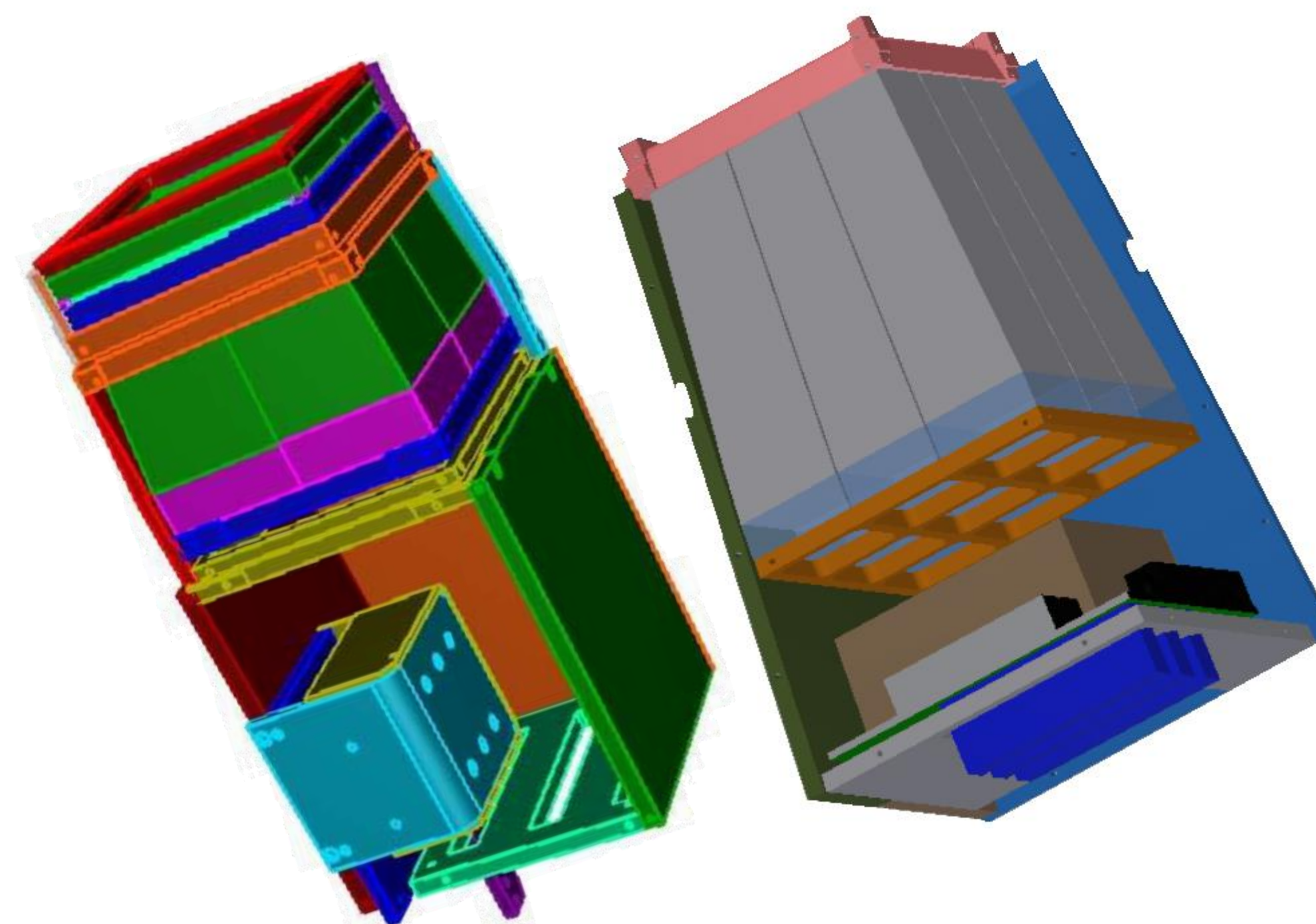
4 cm CsI Stops Proton of 116 MeV



10 cm CsI Stopped Proton of 190 MeV



Old and New Telescope Design



REDESIGNED COMPONENTS

- CsI Crystals:** Longer cesium iodide crystals were used to detect high energy particles.
- Telescope Box:** A new box was designed with fewer parts to reduce machining costs and to accommodate the new crystals.
- Back Plate:** The back plate had to be altered to accommodate the new box design and to allow cooling of the pre-amps from the back.
- Pre-Amp Box:** The new pre-amp box was designed in order to hold 9 pre-amps instead of 4.

ACKNOWLEDGEMENTS:
Juan Manfredi, Sean Sweany



HiRA is funded in part by NSF Major Research Instrumentation (MRI) grant PHY-1102511, and is a collaboration between:



The NSCL is funded in part by Michigan State University and the National Science Foundation.

