

# The neutron long counter NERO for studies of $\beta$ -delayed neutron emission in the r-process

J. Pereira<sup>\*,a,b</sup>, P. Hosmer<sup>a,b,c,1</sup>, G. Lorusso<sup>a,b,c</sup>, P. Santi<sup>a,b,2</sup>, A. Couture<sup>h,i,j</sup>,  
J. Daly<sup>h,i,j</sup>, M. Del Santo<sup>a,b</sup>, T. Elliot<sup>a,b,c</sup>, J. Görres<sup>h,i,j</sup>, C. Herlitzius<sup>d,b</sup>,  
K.-L. Kratz<sup>e,f</sup>, L.O. Lamm<sup>h,i,j</sup>, H.Y. Lee<sup>h,i,j</sup>, F. Montes<sup>a,b</sup>, M. Ouellette<sup>a,b,c</sup>,  
E. Pellegrini<sup>a,b,c</sup>, P. Reeder<sup>k</sup>, H. Schatz<sup>a,b,c</sup>, F. Schertz<sup>d,f</sup>,  
L. Schnorrenberger<sup>a,b,g</sup>, K. Smith<sup>a,b,c</sup>, E. Stech<sup>h,i,j</sup>, E. Strandberg<sup>h,i,j</sup>,  
C. Ugalde<sup>h,i,j</sup>, M. Wiescher<sup>h,i,j</sup>, A. Wöhr<sup>h,i,j</sup>

<sup>a</sup>*National Superconducting Cyclotron Laboratory, Michigan State University,  
East Lansing, Michigan, USA*

<sup>b</sup>*Joint Institute for Nuclear Astrophysics, Michigan State University,  
East Lansing, Michigan, USA*

<sup>c</sup>*Department of Physics and Astronomy, Michigan State University,  
East Lansing, Michigan, USA*

<sup>d</sup>*Institut für Kernchemie, Universität Mainz, Mainz, Germany*

<sup>e</sup>*Max-Planck-Institut für Chemie, Universität Mainz, Mainz, Germany*

<sup>f</sup>*Virtuelles Institut für Struktur der Kerne and Nuklearer Astrophysik, Mainz, Germany*

<sup>g</sup>*Institut für Kernphysik, TU Darmstadt, Darmstadt, Germany*

<sup>h</sup>*Institute of Structure and Nuclear Astrophysics, University of Notre Dame,  
South Bend, Indiana, USA*

<sup>i</sup>*Department of Physics and Astronomy, University of Notre Dame,  
South Bend, Indiana, USA*

<sup>j</sup>*Joint Institute for Nuclear Astrophysics, University of Notre Dame,  
South Bend, Indiana, USA*

<sup>k</sup>*Pacific Northwest National Laboratory, Richland, Washington, USA*

---

## Abstract

The neutron long counter NERO was built at the National Superconducting Cyclotron Laboratory (NSCL), Michigan State University, for measuring  $\beta$ -

---

\*Corresponding author. Tel.: +1 517 9087428; fax: +1 517 3535967

*Email address:* pereira@nscl.msu.edu (J. Pereira)

<sup>1</sup>Present address: Navy Nuclear Power School, Goose Creek, South Carolina, USA

<sup>2</sup>Present address: Los Alamos National Laboratory, MS E540, Los Alamos, New Mexico, USA

delayed neutron-emission probabilities. The detector was designed to work in conjunction with a  $\beta$ -decay implantation station, so that  $\beta$  decays and  $\beta$ -delayed neutrons emitted from implanted nuclei can be measured simultaneously. The high efficiency of about 40%, for the range of energies of interest, along with the small background, are crucial for measuring  $\beta$ -delayed neutron emission branchings for neutron-rich r-process nuclei produced as low intensity fragmentation beams in in-flight separator facilities.

*Key words:* Large neutron counter,  $\beta$ -delayed neutron emission,

Astrophysical r-process, Neutron detection efficiency, Neutron background

*PACS:* 28.20.-v, 28.20.Gd, 29.40.-n, 29.40.Cs, 23.40.-s, 25.40.Ny

---

## 1. Introduction

The emission of  $\beta$ -delayed neutrons by neutron-rich nuclei significantly influences [1] the nucleosynthesis of heavy elements in the rapid (r-) neutron-capture process [2, 3]. This decay mechanism competes with the  $\beta$  decay of r-process nuclei towards the valley of stability and serves as an additional source of neutrons in late stages of the r-process [4]. Measurements of  $\beta$ -delayed neutron emission probabilities ( $P_n$ ) are needed for reliable r-process model calculations, and to test the astrophysical assumptions in various r-process models by comparing their final abundance predictions with observations.

From a nuclear-structure point of view, the  $P_n$  value provides model constraints at low beam intensities where  $\gamma$ -spectroscopy is difficult. The  $P_n$  value probes  $\beta$ -decay strength at excitation energies slightly above the neutron threshold. It therefore provides nuclear structure information comple-

15 mentary to  $\beta$ -decay, which often favors low energy  $\beta$ -decay strength owing  
16 to the larger phase space (see for example [5, 6]).

17 The experimental determination of  $P_n$  requires the measurement of  $\beta$ -  
18 delayed neutrons in coincidence with the  $\beta$  particles emitted from the nu-  
19 cleus of interest. This is particularly challenging for nuclei near or at the  
20 r-process path due to their very low production rates and the relatively short  
21 half-lives—of the order of 10–100 milliseconds. Experiments performed at  
22 ISOL-type facilities have successfully exploited the use of neutron long coun-  
23 ters (NLC) [7] to measure  $P_n$  values of neutron-rich nuclei (see, for instance,  
24 the compilations of Refs. [8, 9]). NLCs generally consist of a series of gas pro-  
25 portional counters embedded into a moderator block used to thermalize the  
26 neutrons prior to their detection. Performance requirements include a high  
27 detection efficiency for neutron energies ranging from a few keV to  $\approx 1$  MeV.  
28 Because the detector does not measure the energy of individual neutrons,  
29 variations of the efficiency as a function of energy have to be minimized as  
30 much as possible as they otherwise can translate into uncertainties in the  
31 measured  $P_n$ . Our goal was to keep detector induced systematic uncertain-  
32 ties well below the 10% level. Measurements at that level of accuracy are  
33 a dramatic improvement over theoretical predictions, and ensure that other  
34 uncertainties dominate astrophysical and nuclear structure models. With  
35 systematic errors at that level, statistical errors will tend to dominate in  
36 practice, as the most interesting isotopes will typically be produced at rather  
37 low rates.

38 We report here the development of NERO, a new NLC at National Super-  
39 conducting Cyclotron Laboratory (NSCL) suitable for use with fast radioac-

40 tive beams produced by in-flight fragmentation. This technique provides  
41 exotic beams without some of the limitations induced by chemistry-based  
42 target-extraction techniques. The short time required to transport, sepa-  
43 rate, and identify the produced fragments, typically less than few hundred  
44 ns, makes it possible to study the very short-lived nuclei in the r-process. The  
45 fragments of interest are implanted in an active catcher that is part of the  
46 NSCL Beta Counting System (BCS). Implantation of a fragment and emit-  
47 ted  $\beta$  particles are detected event-by-event. The correlation of decays with  
48 a previously implanted nucleus requires large area highly pixelated catchers,  
49 typically double-sided silicon strip detectors (DSSDs). The challenge in the  
50 design of NERO was to include a large cylindrical cavity capable of accom-  
51 modating such a system, while still fulfilling the performance requirements  
52 for the detection efficiency. The final design was inspired by existing NLC  
53 detectors such as the Mainz Neutron Detector [10].

## 54 **2. Technical aspects**

### 55 *2.1. Design*

56 The detector system consists of a  $60 \times 60 \times 80$  cm<sup>3</sup> polyethylene matrix  
57 (density 0.93(1) g/cm<sup>3</sup>) with its long symmetry axis aligned with the beam.  
58 Along the beam axis, the matrix has a cylindrical cavity with a diameter of  
59 22.8 cm to accommodate the BCS (see Fig. 1, left).

60 NERO includes three different types of cylindrical proportional counters  
61 manufactured by Reuter-Stokes: filled with <sup>3</sup>He (models RS-P4-0814-207 and  
62 RS-P4-0810-104), and filled with BF<sub>3</sub> (model RS-P1-1620-205) (see Tab. 1  
63 for details). Sixty of these detectors are arranged in three concentric rings

64 around the central symmetry axis, allowing for a nearly  $4\pi$  solid angle cov-  
65 erage around the implantation detector (see Fig. 1, right). The optimum  
66 detector configuration was found using the MCNP code [11] to calculate the  
67 neutron-detection efficiency for different geometries, moderating materials,  
68 and number and arrangement of various types of proportional counters. In-  
69 teractions of neutrons with the different detector materials were calculated,  
70 using the ENDF/B-VI [12] cross-sections in the energy range  $10^{-5}$  eV to  
71 20 MeV. The influence of different environments such as laboratory floor and  
72 wall configurations were investigated but were found to be negligible. Ac-  
73 cording to these calculations, most of the neutrons emitted from the center  
74 of NERO are detected in the innermost ring. Therefore, sixteen of the more  
75 compact and efficient  $^3\text{He}$  gas-filled proportional counters are mounted in  
76 the innermost ring at a radius of 13.6 cm. For the middle and outer rings  
77 at radii of 19.2 cm and 24.8 cm we use twenty and twenty-four  $\text{BF}_3$  propor-  
78 tional counters, respectively. The  $\text{BF}_3$  counters are longer allowing one to  
79 cover more solid angle, and their efficiency just compensates the decreasing  
80 efficiency of the inner ring with increasing neutron energy.

81 To facilitate transportation and assembly, the polyethylene block is di-  
82 vided into an upper and lower half, and each half is subdivided in six equal  
83 parts along the longest symmetry axis. The twelve pieces are held together  
84 with eight stainless steel bolts.

## 85 *2.2. Electronics*

86 The NERO readout channels are grouped in 4 quadrants with 15 chan-  
87 nels each (4  $^3\text{He}$  counters and 11  $\text{BF}_3$  counters). Figure 2 shows a schematic  
88 diagram of the NERO electronics for one quadrant. The proportional coun-

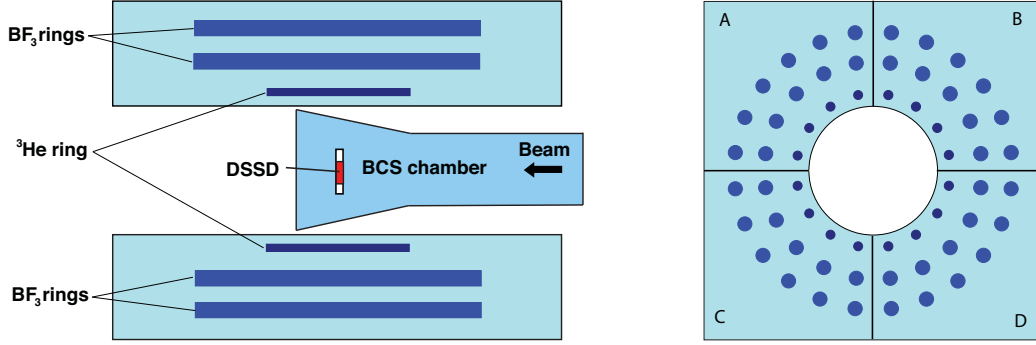


Figure 1: Schematic drawings of the NERO detector. Left: Side view showing the BCS chamber located inside of NERO with the DSSD at the central position. Right: backside showing the cylindrical cavity to house the BCS and the three concentric rings of gas-filled proportional counters. The labels A, B, C and D designate the four quadrants.

Table 1: Technical specifications of the NERO gas-filled proportional counters. (a) and (b) refer to the <sup>3</sup>He detector models RS-P4-0810-104 and RS-P4-0814-207, respectively.

Detector	Active Length (cm)	Radius (cm)	Nominal Pressure (atm)	Gas Composition	High Voltage (+V)
<sup>3</sup> He (a)	25.0(2)	1.3(2)	10.2	100% <sup>3</sup> He	1350
<sup>3</sup> He (b)	35.6(2)	1.3(2)	4.0	100% <sup>3</sup> He	1100
BF <sub>3</sub>	50.8(1)	2.5(2)	1.2	>96% <sup>10</sup> B	600

89 ters detect the charged particles produced in the exothermic neutron-capture  
90 reaction <sup>3</sup>He(*n*,*p*) or <sup>10</sup>B(*n*,*α*), respectively. Their signals feed 16-channel  
91 preamplifiers built at NSCL using Cremat CR-101D miniature charge-sensitive

92 preamp chips. The pre-amplified signals are sent into four 16-channel shaper  
93 and discriminator modules, designed at Washington University, St. Louis,  
94 and manufactured by Pico Systems [13]. These modules integrate indepen-  
95 dent shaping and discriminating circuits sharing the same input. Shaping  
96 times and pole-zero cancelation are adjusted for each channel by properly se-  
97 lecting the capacitances. The gain and threshold levels of the shaper/discriminator  
98 are adjusted via computer control over the CAMAC bus.

99 The logic signals from the discriminator are recorded in scalers and in  
100 a 64-channel multi-hit (VME) TDC that is common for all quadrants. The  
101 TDC was programmed to work in start-gate mode, in which a gate signal,  
102 generated by a  $\beta$  decay detected in the BCS, enables the module to accept  
103 multiple stop signals in each channel from any of the sixty gas counters. The  
104 duration of this gate ( $\tau=200 \mu\text{s}$ ) was chosen to account for the time needed  
105 to moderate and detect the neutrons (see Sec. 3.1). The  $P_n$  value of a given  
106 nucleus is extracted from the number of stops-signals registered in the TDC  
107 (i.e., neutrons correlated with  $\beta$  decays) relative to the number of  $\beta$  decays  
108 detected in the BCS.

109 The shaper outputs are connected to 32-channel (VME) ADC cards. The  
110 pulse height spectra recorded by the ADCs are used to set the thresholds  
111 and gains of the shaper/discriminator units, and to monitor any background  
112 or gain variation during the course of an experiment. Figure 3 shows typical  
113 ADC spectra for  $^3\text{He}$  and  $\text{BF}_3$  gas counters recorded under different condi-  
114 tions. The spectra show the typical wall-effect: The location of the peak  
115 at high amplitudes marks the Q value of the neutron-capture reaction, i.e.,  
116  $^3\text{He}(n,p)t$  and  $^{10}\text{B}(n,\alpha)^7\text{Li}$  for the  $^3\text{He}$  and  $\text{BF}_3$  gas counters, respectively;

117 the plateau or low energy tail at low amplitudes arises from events where  
 118 reaction products hit the detector wall preventing the complete deposition of  
 119 their energy in the counter gas. Thresholds are set below these low amplitude  
 120 events and just above the tail of the prominent low energy peak generated  
 121 by electronic noise and background  $\gamma$  radiation. Note that unlike the TDC,  
 122 the ADCs only register one neutron per BCS trigger.

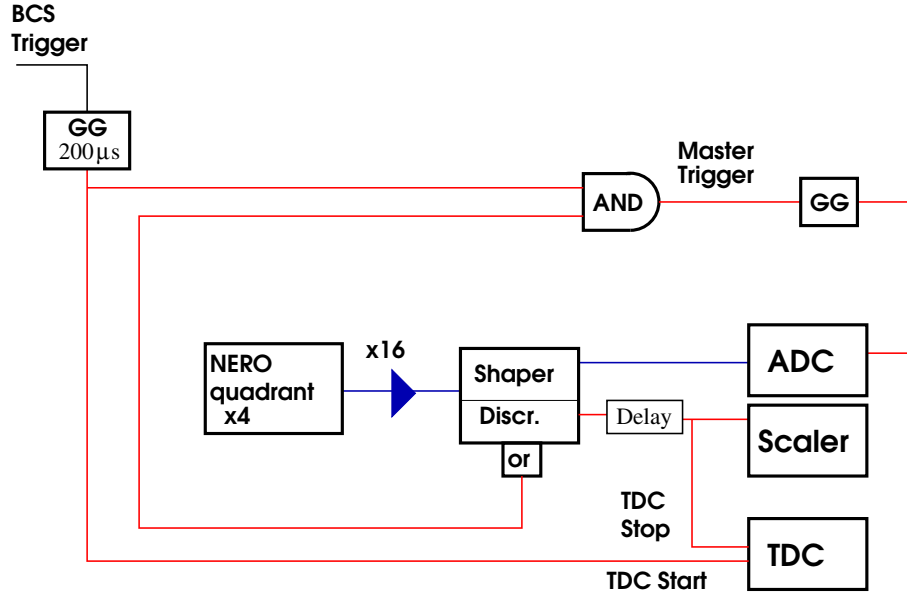


Figure 2: NERO electronic diagram (see text for details). For clarity only one NERO quadrant is shown (GG stands for Gate Generator).

### 123 3. Detector performance

#### 124 3.1. Moderation time

125 In order to determine the optimal duration of the TDC gate  $\tau$ , we mea-  
 126 sured the time needed by the neutrons to slow down in the polyethylene



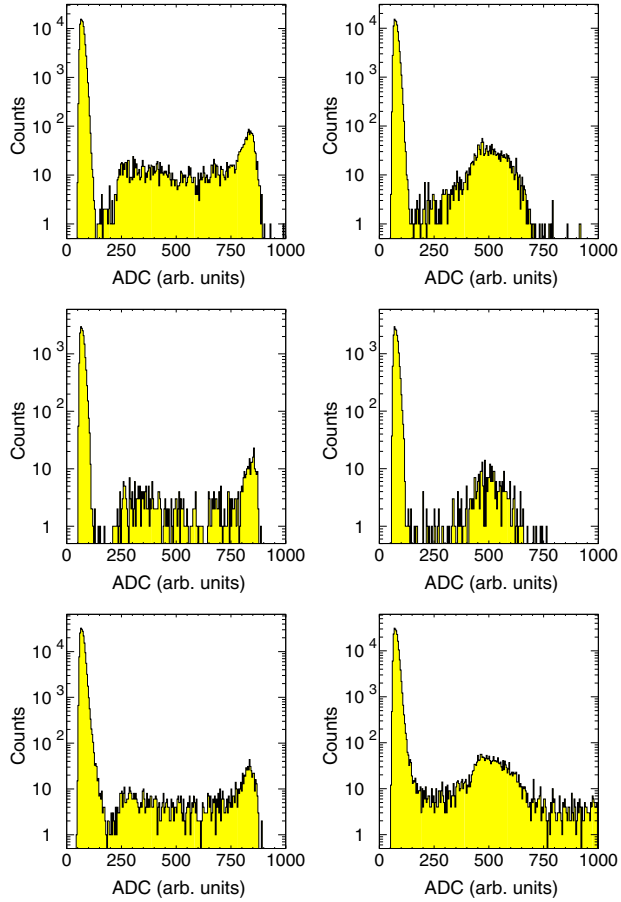


Figure 3: NERO ADC spectra for one of the  $^3\text{He}$  (left panels) and  $\text{BF}_3$  (right panel) gas counters. The top panels were recorded during 5 minutes, using a  $^{252}\text{Cf}$  neutron source at the center of NERO. The middle panels correspond to a one-hour measurement with nuclei produced in the reaction  $^{136}\text{Xe}$  (150 MeV/u)+Be that were implanted in the BCS and included  $\beta$  delayed neutron emitters. The data displayed in the bottom panels were recorded during a 12-hour background measurement without beam on target using NERO as trigger.

127 moderator before their detection ( $\tau_n$ ). A  $^{252}\text{Cf}$  source was located at the  
 128 center of NERO facing a NaI scintillator at a distance of 5 cm. Neutrons

129 and  $\gamma$  rays were emitted in coincidence from the fragments produced in the  
 130 spontaneous fission of  $^{252}\text{Cf}$ . The scintillator was used to detect the  $\gamma$  rays,  
 131 which provided the external trigger of the NERO electronics (replacing the  
 132 BCS trigger shown in Fig. 2). The time difference between the detection of  
 133 a  $\gamma$  ray and the subsequent moderated neutrons recorded in the multi-hit  
 134 TDC provided  $\tau_n$ . The distribution of  $\tau_n$  is shown in Fig. 4 (left) for the  
 135 innermost, intermediate, and outer rings of proportional counters, as well as  
 136 for the whole detector.

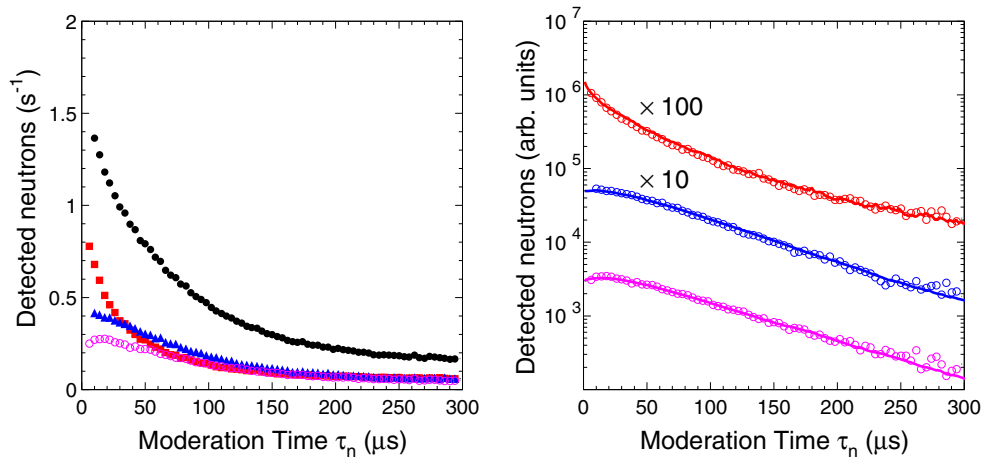


Figure 4: (Color online). Left: Measured background-subtracted moderation time distributions for a  $^{252}\text{Cf}$  source for the innermost ring (squares), the intermediate ring (triangles), the outer ring (empty circles) and for the entire detector (solid circles). Right: Moderation-time distribution for each ring (empty circles), compared with results obtained with MCNP (solid lines). For easy comparison, the measured and calculated distributions for the first and second rings are scaled by 100 and 10, respectively. Note that the error bars of the experimental data are smaller than the symbol sizes.

137 The largest differences between the three rings are found at the shortest

138 times, when most of the neutrons emitted from the center of NERO reach  
139 the innermost ring. At late times, the neutrons are more uniformly dis-  
140 tributed over the whole moderator, and the three rings have similar detection  
141 rates. For each ring, the excellent agreement of the measured, background-  
142 subtracted moderation-time distributions with MCNP simulations is shown  
143 in Fig. 4. Between  $50 \mu\text{s}$  and  $300 \mu\text{s}$  the time distributions can be approxi-  
144 mated with exponential functions. The corresponding measured and calcu-  
145 lated moderation time scales are  $43 \mu\text{s}$  and  $41 \mu\text{s}$  for the first ring,  $51 \mu\text{s}$  and  
146  $52 \mu\text{s}$  for the second ring, and  $55 \mu\text{s}$  and  $59 \mu\text{s}$  for the third ring, respectively.

147 From Fig. 4, we find that  $94.3(1)\%$  of the neutrons are detected within  
148  $\tau_n \leq 200 \mu\text{s}$ . The energies  $E_n$  of the neutrons emitted in the spontaneous  
149 fission of  $^{252}\text{Cf}$  are typically described by a Maxwell-Boltzmann distribution  
150 function with an effective temperature of  $kT=1.42 \text{ MeV}$  [14], an average  
151 neutron energy of  $2.1 \text{ MeV}$  and a smooth tail at higher energies that extends  
152 up to  $9 \text{ MeV}$ . These are higher energies than typically expected for  $\beta$ -delayed  
153 neutrons (see discussion at the end of Sec. 3.2.4). Since the moderation time  
154 increases with neutron energy, we expect that more than  $94.3(1)\%$  of  $\beta$ -  
155 delayed neutrons are detected within  $200 \mu\text{s}$ . We therefore chose a TDC gate  
156 of  $\tau=200 \mu\text{s}$ .

### 157 3.2. Efficiency

158 In order to characterize the NERO efficiency and its energy dependence,  
159 different types of measurements were performed using a  $^{252}\text{Cf}$  neutron source  
160 of known activity and neutrons produced in resonant and non-resonant reac-  
161 tions at the Institute for Structure and Nuclear Astrophysics (ISNAP) at the  
162 University of Notre Dame. The results obtained from these reactions were

163 used to constrain the energy dependence of the NERO efficiency.

### 164 *3.2.1. Measurement of NERO efficiency with a $^{252}\text{Cf}$ neutron source*

165 Before and after the calibration measurements performed at the Univer-  
166 sity of Notre Dame, the NERO efficiency was measured with a  $1.251(5) \mu\text{Ci}$   
167  $^{252}\text{Cf}$  calibration source with an active diameter of 5 mm (neutron branching  
168 11.6% and half-life 2.689 years). Additional contributions to the total neu-  
169 tron rate from contaminants were estimated. Besides  $^{252}\text{Cf}$ , there are small  
170 impurities of  $^{249-251}\text{Cf}$  and  $^{254}\text{Cf}$ .  $^{249}\text{Cf}$  and  $^{251}\text{Cf}$  have a negligibly small  
171 spontaneous fission branch, whereas the present amount of  $^{254}\text{Cf}$  was very  
172 small due to its short half-life. Consequently, only the  $^{250}\text{Cf}$  and  $^{252}\text{Cf}$  iso-  
173 topes had to be considered. At the time of the measurement, 3.6% of the total  
174 activity of the source was due to  $^{250}\text{Cf}$ , whose contribution to the neutron  
175 activity was negligible due to its very low neutron branching of 0.296%. The  
176 alpha decay of  $^{252}\text{Cf}$  produces  $^{248}\text{Cm}$ , which undergoes spontaneous fission  
177 accompanied with neutron emission, with a branching ratio of 8.39%. The  
178 very long half-life of this radioisotope ( $3.48 \times 10^5$  years) made its contribution  
179 to the total neutron rate negligible.

180 The number of detected neutrons was recorded with scalers and the multi-  
181 hit TDC described in Sec. 2.2. We verified that data processing dead time  
182 is negligible up to 50 KHz, well above the activity of the source, using a  
183 random pulser. Similarly, the  $2 \mu\text{s}$  dead-time in the proportional counters  
184 was negligible. Taking the ratio of the number of neutrons recorded with  
185 NERO to the number of neutrons emitted by the source, calculated from  
186 the known source activity and neutron branching, we obtained a neutron  
187 detection efficiency of  $31.7(2)\%$ . This value also serves as a reference point

188 to verify the NERO efficiency before, during, and after experiments.

189 *3.2.2. Measurement of NERO efficiency with resonant reactions*

190 The two resonant reactions used to study the NERO efficiency were  
 191  $^{13}\text{C}(\alpha,n)^{16}\text{O}$  [15–17] and  $^{11}\text{B}(\alpha,n)^{14}\text{N}$  [18, 19]. The ISNAP KN Van de Graff  
 192 accelerated a beam of  $\alpha$  particles impinging onto  $^{13}\text{C}$  and  $^{11}\text{B}$  targets of  
 193  $14(2) \mu\text{g}/\text{cm}^2$  and  $12_{-2}^{+4} \mu\text{g}/\text{cm}^2$  thickness, respectively, located at the cen-  
 194 ter of the NERO symmetry axis. The rate of incident  $\alpha$  particles ( $I_\alpha$ ) was  
 195 monitored with an isolated electron-suppressed plate behind the target. A  
 196 total of two resonances for the reaction  $^{13}\text{C}(\alpha,n)^{16}\text{O}$ , and one for the reac-  
 197 tion  $^{11}\text{B}(\alpha,n)^{14}\text{N}$  were used (see Table 2). Each resonance was completely  
 198 mapped around its peak energy  $E_R$  by detecting the number of neutrons  
 199 as a function of  $\alpha$ -beam energy  $E_\alpha$ . A linear background function was fit  
 200 underneath the resonance curve and used to subtract non-resonant contri-  
 201 butions and background neutrons. The efficiency of NERO was determined  
 202 as the ratio of the number of detected neutrons to the number of neutrons  
 203  $N_n$  produced in the resonant reaction. Since the resonances considered here  
 204 fulfill  $\Gamma_R \ll \Delta E = E_i - E_f$  and  $\Gamma_R \ll E_R$  (see Table 3), one can use  
 205 the thick-target narrow-resonance approximation and calculate  $N_n$  (see for  
 206 example [20]) using

$$N_n = \frac{I_\alpha t \pi^2 \hbar^2 (\omega\gamma)_R N_A \rho}{\mu A E_R} \left( \frac{dE}{dz} \right)^{-1}. \quad (1)$$

207 where  $(dE/dz)$  is the stopping power of the  $\alpha$  particle in the target material,  
 208 calculated with the SRIM-2000 code [21] in the center-of-mass frame;  $t$  is  
 209 the duration of the measurement;  $A$  is the mole mass of the target;  $\rho$  is the  
 210 target mass density;  $\mu$  is the reduced mass of the system and  $(\omega\gamma)_R$  is the

211 resonance strength. The resonance parameters used in the calculations were  
 212 taken from Refs. [15–19]. They are summarized in Table 2.

213 The results of the efficiency measurements are shown in Table 4 for the  
 214 three selected resonances, along with their corresponding evaluated average  
 215 neutron energies in the laboratory frame  $\langle E_n \rangle$ . The calculation of this latter  
 216 quantity is described in Sec. 3.2.4.

Table 2: Properties of the three selected resonances: resonance energy in the laboratory frame  $E_\alpha$ , resonance width  $\Gamma_R$ , strength  $(\omega\gamma)_R$ , excitation energy  $E_x$ , and spin and parity  $J^\pi$ .

Reaction	$E_\alpha$ (MeV)	$\Gamma_R$ (keV)	$(\omega\gamma)_R$ (eV)	$E_x$ (keV)	$J^\pi$
$^{13}\text{C}(\alpha,n)^{16}\text{O}$	1.053	1.5(2)	11.9(6)	7165	5/2 <sup>-</sup>
$^{13}\text{C}(\alpha,n)^{16}\text{O}$	1.585	≤1	10.8(5)	7576(2)	7/2 <sup>-</sup>
$^{11}\text{B}(\alpha,n)^{14}\text{N}$	0.606	$2.5(5)\times 10^{-3}$	0.175(10)	11436	7/2 <sup>-</sup>

217 *3.2.3. Measurement of NERO efficiency with non-resonant reactions*

218 Additional measurements of the NERO efficiency were performed at IS-  
 219 NAP using neutrons produced in the  $^{51}\text{V}(p,n)^{51}\text{Cr}$  reaction [22] at three  
 220 different energies. This reaction has been used in the past for neutron de-  
 221 tector calibrations [16, 18]. Here, a proton beam was accelerated at the KN  
 222 accelerator and impinged onto a  $^{51}\text{V}$  target mounted in the center of NERO.  
 223 Three incident proton energies of 1.8 MeV, 2.14 MeV and 2.27 MeV were  
 224 chosen from regions of the excitation function with no individual resonances,

Table 3: Validation of the thick-target, narrow-resonance approximation.  $d\rho$  is the target thickness and  $\Delta E$  is the energy loss in the target at the resonance energy.

Reaction	$E_\alpha$ (MeV)	$d\rho$ (mg/cm <sup>2</sup> )	$\Delta E$ (keV)	$\Gamma_R$ (keV)
$^{13}\text{C}(\alpha,n)^{16}\text{O}$	1.053	0.014	18	1.5(2)
$^{13}\text{C}(\alpha,n)^{16}\text{O}$	1.585	0.014	16	$\leq 1$
$^{11}\text{B}(\alpha,n)^{14}\text{N}$	0.606	0.012	18	$2.5(5)\times 10^{-3}$

225 using three targets with a thickness of 32  $\mu\text{g}/\text{cm}^2$ .

226 To determine the number of  $^{51}\text{V}(p,n)^{51}\text{Cr}$  reactions that have occurred  
227 during a measurement one can take advantage of the fact that for every  
228  $^{51}\text{V}(p,n)^{51}\text{Cr}$  reaction, a radioactive  $^{51}\text{Cr}$  is created with a half-life of 27.7025(24)  
229 days. The electron-capture decay of  $^{51}\text{Cr}$  is followed by the emission of sev-  
230 eral X-rays [23] and a 320.0824(4) keV  $\gamma$  ray [24] with a branching ratio of  
231 9.91(1)% [23]. The number of neutrons  $N_n$  produced in the reaction can then  
232 be simply determined from the activity of the target after irradiation. The  
233 number of 320.1 keV  $\gamma$  rays emitted was measured offline in a lead-shielded  
234 setup, where the irradiated target was mounted in a plastic holder facing a  
235 HPGe  $\gamma$  detector. Decay losses during the irradiation, transport, and offline  
236 counting were negligible. The HPGe efficiency at 320.1 keV was found to  
237 be 0.76(4)% using a  $^{133}\text{Ba}$  calibration source, and by interpolating the effi-  
238 ciencies measured for the two  $\gamma$  rays emitted at 302 keV and 356 keV. The  
239 deduced NERO neutron efficiencies are listed in Table 4 for the three different

240 proton energies. The systematic error is dominated by the 5% uncertainty  
 241 in the activity of the  $\gamma$ -ray calibration source.

Table 4: Laboratory frame projectile energy  $E_{proj}$ , average neutron energy  $\langle E_n \rangle$ , and width of the neutron energy distribution  $\Delta E_n$  with corresponding measured neutron detection efficiency  $\epsilon_n$ . The last three columns list efficiencies for isotropic sources emitting neutrons at the energy  $\langle E_n \rangle$ .  $\epsilon_n$  “isotr” is derived from the experimental value  $\epsilon_n$  using corrections calculated with MCNP,  $\epsilon_n$  “MCNP” is the calculated efficiency, and  $\epsilon_n$  “MCNP scaled” is the calculated efficiency scaled to obtain a best fit to the experimental data.

Reaction	$E_{proj}$	$\langle E_n \rangle$	$\Delta E_n$	$\epsilon_n$	$\epsilon_n$	$\epsilon_n$	$\epsilon_n$
					isotr.	MCNP	MCNP scaled
	(MeV)	(MeV)	(MeV)	(%)	(%)	(%)	(%)
$^{11}\text{B}(\alpha,n)$	0.606	0.56	0.12	33(2)	38(2)	42	37
$^{13}\text{C}(\alpha,n)$	1.053	2.8	0.31	24(1)	30(1)	29	26
$^{13}\text{C}(\alpha,n)$	1.585	3.2	0.41	27(1)	33(1)	29	24
$^{51}\text{V}(p,n)$	1.80	0.23	0.014	39(2)	36(2)	44	39
$^{51}\text{V}(p,n)$	2.14	0.55	0.024	34(2)	32(2)	42	37
$^{51}\text{V}(p,n)$	2.27	0.68	0.028	34(2)	34(2)	41	36

#### 242 3.2.4. Results and discussion

243 In order to evaluate the energy-dependence of the efficiency, the en-  
 244 ergy spectrum of the emitted neutrons needs to be known for each reaction



245 used. These neutrons are, for our purposes, essentially mono-energetic in  
 246 the center-of-mass frame. The center-of-mass energies  $\hat{E}_n$  can be calculated  
 247 from the known reaction Q-values. However, the corresponding laboratory  
 248 frame neutron energy  $E_n$  depends on the center-of-mass polar angle  $\hat{\theta}$  of the  
 249 emitted neutron with respect to the beam axis. This leads to a broadening  
 250 of the neutron energy distribution, with an average neutron energy  $\langle E_n \rangle$  of

$$\langle E_n \rangle = \int_{-1}^{+1} E_n(x)W(x)dx, \quad (2)$$

251 where  $x = \cos \hat{\theta}$ . The angular-correlation functions  $W(x)$  for each of the three  
 252 resonances used here were calculated as described in [25] and are shown in  
 253 Fig. 5. In the case of the  $^{11}\text{B}(\alpha, n)^{14}\text{N}$  reaction, the coupling of the neutron  
 254 spin with the ground-state of  $^{14}\text{N}$  leads to two possible values of the final spin.  
 255 Variations of the results obtained using the two possible  $W(x)$  functions were  
 256 included in the final uncertainty. For  $^{51}\text{V}(p, n)^{51}\text{Cr}$  we have chosen energies  
 257 where the excitation function shows non resonant behavior. This justifies  
 258 the assumption of isotropic neutron emission in the center-of-mass frame, or,  
 259 equivalently,  $W(x) = 1/2$ .

260 The average neutron energies calculated in the laboratory frame  $\langle E_n \rangle$  to-  
 261 gether with the width of the energy distribution  $\Delta E_n$  are shown in Table 4 for  
 262 each reaction, along with the corresponding measured efficiencies  $\epsilon_n$ . In the  
 263 case of  $^{51}\text{V}(p, n)^{51}\text{Cr}$   $\Delta E_n$  also includes a small contribution from the energy  
 264 loss and straggling in the target. In order to determine the efficiencies for neu-  
 265 trons emitted isotropically with a given energy from the DSSD catcher in the  
 266 BCS, our measured  $\epsilon_n$  need to be corrected for the angular distribution and  
 267 energy range of the neutrons (see e.g. Ref. [26] for the reaction  $^{13}\text{C}(\alpha, n)^{16}\text{O}$ ).  
 268 This was done using MCNP simulations: First, the NERO efficiencies were

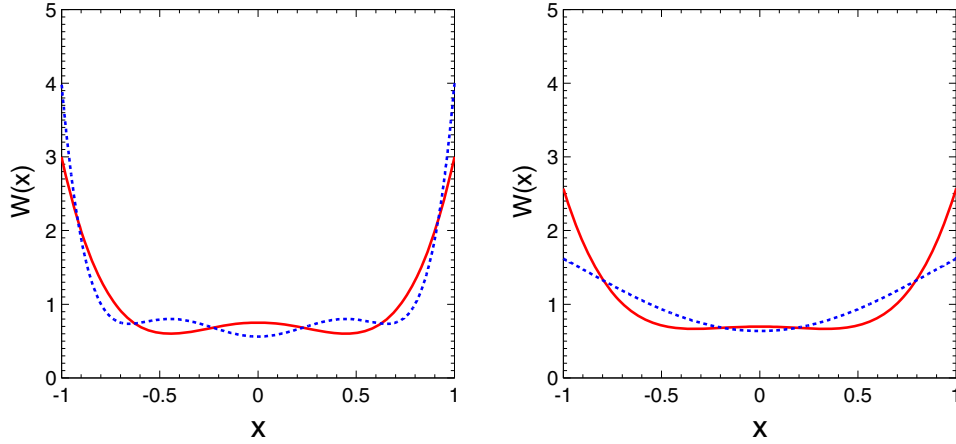


Figure 5: Left: Angular correlation  $W(x)$  calculated in the center-of-mass frame for the 7.165 MeV,  $5/2^-$  (solid line) and 7.576 MeV,  $7/2^-$  (dashed line) resonances in the reaction  $^{13}\text{C}(\alpha,n)^{16}\text{O}$ . Right: Angular correlation  $W(x)$  calculated in the center-of-mass frame for  $J_f=1/2^+$  (solid line) and  $J_f=3/2^+$  (dashed line) of the 11.436 MeV,  $7/2^-$  resonance in the resonant reaction  $^{11}\text{B}(\alpha,n)^{14}\text{N}$ .

269 calculated with MCNP at the energies  $\langle E_n \rangle$  of Table 4, assuming an isotropic  
 270 mono-energetic neutron source located in the center of the detector. A sec-  
 271 ond calculation was then performed, using the calculated laboratory frame  
 272 angular and energy distributions of the neutrons. The MCNP-calculated  
 273 anisotropic-to-isotropic efficiency ratios were then used as a correction factor  
 274 to translate the measured efficiencies into efficiencies for isotropic emission  
 275 at a single energy  $\langle E_n \rangle$ . In Table 4, we show the measured efficiencies for  
 276 the different reactions ( $\epsilon_n$ ) and the corresponding corrected values ( $\epsilon_n$  isotr.).  
 277 The strongest correction of about 15% arises mainly from the angular cor-  
 278 relation  $W(x)$  in the resonant reactions. The isotropic efficiencies are shown

279 in Fig. 6 for the whole detector (left), and for each ring separately (right).

280 The experimentally-determined efficiencies covered a range of energies  
281 from about 0.2 MeV to 3 MeV. In order to extrapolate the results to lower  
282 energies, we combined the experimental values ( $\epsilon_n$  isotr.) with a MCNP  
283 calculation of the efficiency as a function of energy (see dotted line in Fig. 6,  
284 left). Despite a global absolute overestimation of about 5%, the calculated  
285  $\epsilon_n(E_n)$  function follows very well the energy dependence obtained from the  
286 measured data.

287 As shown in Fig. 6 (right), the efficiencies calculated independently for  
288 each ring follow reasonably well the energy trends of the experimental data.  
289 For all the reactions investigated, MCNP reproduces the efficiency of the  
290 innermost ring, which is the most efficient of the three. The agreement is  
291 somewhat worse for the other rings at the lowest energies. In particular, the  
292 calculations overestimate the efficiency of the middle ring by about 3% at  
293 energies below 700 keV. We investigated the possibility that this discrepancy  
294 could be related to the type of detectors used. Several test measurements  
295 of the efficiency were performed with a  $^{252}\text{Cf}$  source, using one single  $^3\text{He}$   
296 proportional counter placed in the first, second and third ring, using the  $^{252}\text{Cf}$   
297 source. A comparison of the calculated efficiencies with the values measured  
298 under these conditions was consistent with the results shown in Fig. 6 (right).  
299 The overall 5% absolute overestimation of the efficiency by the calculations  
300 could not be attributed to the uncertainty in the polyethylene density. First  
301 of all, variations of the density modified the calculated efficiencies for the  
302 second and third ring in the opposite direction of the first ring. Secondly,  
303 when the variations in density were limited to the uncertainties provided by

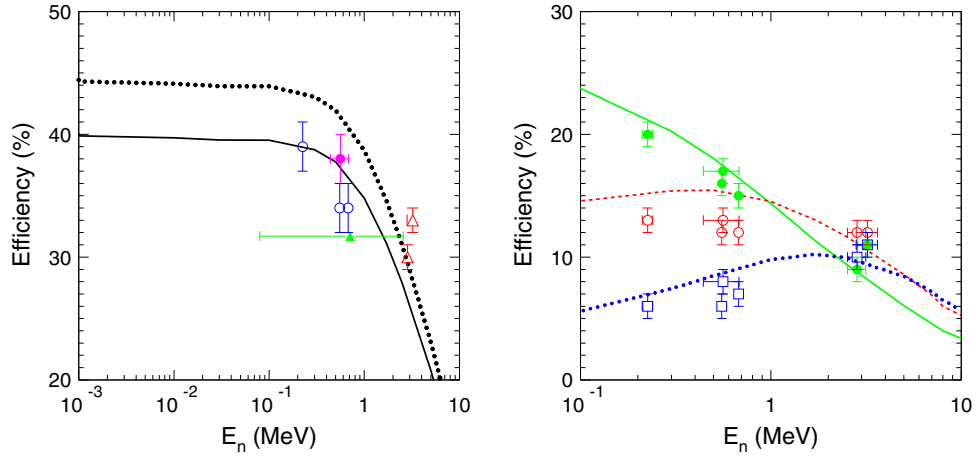


Figure 6: (Color online). Left: MCNP-calculated total efficiency as a function of the neutron energy  $E_n$  scaled to the experimental data (solid line) and unscaled (dotted line), compared with measurements for the reactions  $^{11}\text{B}(\alpha,n)$  (solid circle),  $^{13}\text{C}(\alpha,n)$  (empty triangles),  $^{51}\text{V}(p,n)$  (empty circles), and with the  $^{252}\text{Cf}$  neutron source (solid square). The energy width of the  $^{252}\text{Cf}$  measurement was calculated according to the shortest-interval criterion defined by W. Brüchle for asymmetric distributions [27]. Right: MCNP-calculated efficiencies as a function of neutron energy  $E_n$  for the innermost ring (solid line), intermediate ring (dashed line) and outer ring (dotted line), compared to the measured values for the first (solid circles), second (empty circles) and third (empty squares) rings. Note the different scales of the two figures.

304 the supplier, no differences in the calculated results were observed.

305 The good agreement of MCNP with experimental data observed in Fig. 6  
 306 (right) for the first NERO ring with a thinner moderator layer, and the small  
 307 discrepancies found for the second and third rings with thicker moderator  
 308 layers points to a limitation of MCNP to accurately calculate the scattering  
 309 process of the neutrons in the moderator material. One possibility would be

310 molecular vibrational and rotational excitation modes in the moderator ma-  
 311 terial. Whereas this problem would be hardly observable in detectors with  
 312 thin moderators (e.g. Ref. [28]), it would become more severe for thicker  
 313 moderators. Interestingly, similar conclusions were drawn when comparing  
 314 MCNP calculations with neutron-flux measurements performed with thick  
 315 neutron detectors [29, 30]. In order to compensate for these model defi-  
 316 ciencies we scaled the calculated efficiencies for each ring independently to  
 317 better match the experimental data (see scaled efficiency in Table 4). The  
 318 new scaled efficiency-curve (solid line in Fig. 6, left) can thus be used to  
 319 extrapolate the efficiency to energies below 200 keV.

320 It is worth noting that the relevant neutron energy range for  $\beta$ -delayed  
 321 neutron emission in r-process nuclei is a few hundred keV. As an example, for  
 322 the r-process nuclei around  $A \sim 100 - 130$ , spectroscopic studies of  $\beta$ -delayed  
 323 neutron emitters [31–33] showed that  $\hat{E}_n$  is typically much lower than  $Q_\beta - S_n$ .  
 324 Neutron energies in the laboratory frame  $E_n$  were found to be 199 keV for  
 325  $^{87}\text{Br}$ , 450 keV for  $^{98}\text{Rb}$ , and 579 keV for  $^{137}\text{I}$ . This result was further supported  
 326 by the measured average neutron energies of fission fragments from  $^{235}\text{U}$   
 327 ( $\langle E_n \rangle = 575$  keV) and  $^{239}\text{Pu}$  ( $\langle E_n \rangle = 525$  keV), where, in addition, very few  
 328 neutrons were found at  $E_n \gtrsim 800$  keV [31, 34]. According to these authors,  
 329 the reason for the “compressed”  $E_n$  spectra is the preferred population of  
 330 the lowest excited states in the final nuclei [33]. Our experimental efficiency  
 331 calibration therefore covers the most critical energy range, and the condition  
 332 of an energy independent efficiency for  $\beta$ -delayed neutrons is well fulfilled.  
 333 As an example,  $\epsilon_n$  shown in Fig. 6 (left) shows a relative variation of about  
 334  $\pm 5\%$  for energies below 800 keV. This variation will contribute to the final

335 uncertainty of the measured  $P_n$ . This uncertainty can be reduced if the  
336 neutron energies  $E_n$  can be constrained from experiment or theory.

### 337 3.3. Background

338 One limitation for the measurement of  $P_n$ , particularly for very exotic  
339 nuclei, is the neutron background rate ( $B_n$ ). Its estimation requires to dis-  
340 tinguish two different origins. First, there is the “intrinsic” background as-  
341 sociated with the electronics of the NERO detector and its sensitivity to the  
342 neutrons present in the environment (mainly cosmic rays). Secondly, dur-  
343 ing the course of an experiment, there are beam-induced neutrons produced  
344 by nuclear reactions unrelated to the  $\beta$ -delayed neutron emission of inter-  
345 est. During experiments, neutron background rates measured with NERO  
346 in self-trigger mode can vary within about 5–10  $s^{-1}$  depending on whether  
347 or not the beam is on target. As will be discussed later, the impact of these  
348 background rates is dramatically reduced when the neutrons are measured  
349 in coincidence with  $\beta$  decays.

350 Analysis of the ring-counting ratios for background runs (self-trigger mode)  
351 and production runs (external trigger mode) support the idea of an external  
352 and a beam-induced background source. As shown in Fig. 7 (left), measure-  
353 ments performed with NERO in external trigger mode (i.e. from  $\beta$  decays in  
354 the BCS) with  $\beta$ -delayed neutron emitters showed that the neutron count-  
355 ing rates were higher for the innermost ring and systematically decreased for  
356 the outer rings, in agreement with the results shown in Fig. 6 (right). On  
357 the other hand, background runs with beam off showed the opposite trend,  
358 with high rates in the outer ring, gradually decreasing for the inner ones  
359 (Fig. 7, center). This result suggests that these runs were mainly affected by

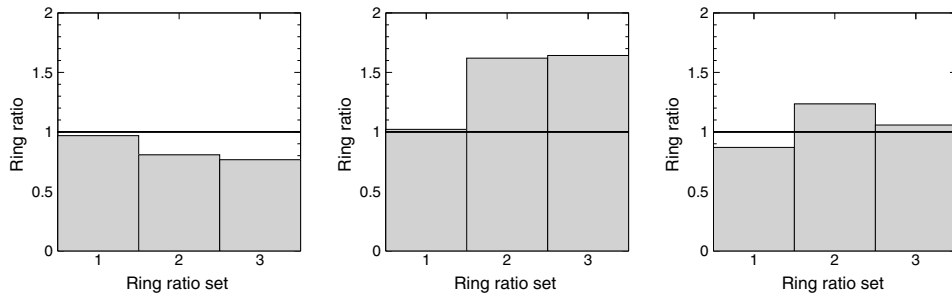


Figure 7: Ratio of neutrons detected with different NERO rings for three different runs: production with  $\beta$ -delayed neutron emitters (left), background with beam off (center), and background with beam on (right). Histogram bin numbers 1, 2 and 3 correspond to ring ratios  $R_2/R_1$ ,  $R_3/R_2$ , and  $R_3/R_1$ , where  $R_{1-3}$  are the innermost, intermediate and external rings. Statistical errors are negligible.

360 an external source of background neutrons, most probably related to cosmic  
 361 rays. Finally, background runs with beam on target showed an intermediate  
 362 situation that could be explained as arising from a combination of external  
 363 and internal sources (Fig. 7, right). Energy spectra obtained for background  
 364 runs with the ADCs show the wall-effect shape expected for neutrons (see  
 365 lower spectra in Fig. 3). Electronic and  $\gamma$ -ray contributions are largely below  
 366 the discriminator thresholds.

#### 367 4. Measurement of $P_n$

368 The NERO detector, together with the BCS, has been employed in nu-  
 369 merous r-process motivated experiments performed at NSCL [5, 6, 35–37].  
 370 The exotic nuclei of interest are implanted in a  $40 \times 40$ -pixel DSSD in the  
 371 BCS.  $\beta$ -decays are also detected in the DSSD and can be position-correlated  
 372 to previously implanted ions during a maximum correlation time  $t_c$ .  $P_n$  values

373 were determined using the number of neutrons  $N_{\beta n}$  detected in coincidence  
 374 with an implantation-correlated  $\beta$ -decay event, according to the equation:

$$P_n = \frac{N_{\beta n} - B_{\beta n} - N_{\beta\beta n}}{\epsilon_n N_\beta}, \quad (3)$$

375 where  $B_{\beta n}$  is the number of background coincidences between  $\beta$ -like events  
 376 (including real  $\beta$  decays and background from the BCS) and neutrons, and  $N_\beta$   
 377 is the number of  $\beta$ -decaying mother nuclei.  $N_{\beta\beta n}$  is the number of detected  
 378  $\beta$ -delayed neutrons from the daughter nuclei and needs to be subtracted from  
 379  $N_{\beta n}$ . For the nuclear species analyzed in Refs. [5, 6, 35–37],  $\beta$ -neutron coin-  
 380 cidences associated with descendant nuclei other than the  $\beta$ -decay daughter  
 381 were negligible. In this case, using the Batemann equations [38], it is possible  
 382 to write explicitly the value of  $N_{\beta\beta n}$  as:

$$N_{\beta\beta n} = (1 - P_n)C, \quad (4)$$

383 where  $C$  is a constant given by:

$$C = \frac{\lambda_2 P_{nn} N_\beta \epsilon_n}{\lambda_2 - \lambda_1} \left[ 1 - e^{-\lambda_1 t_c} - \frac{\lambda_1}{\lambda_2} (1 - e^{-\lambda_2 t_c}) \right]. \quad (5)$$

384 In this equation,  $P_{nn}$  is the neutron-emission probability of the daughter  
 385 nucleus, and  $\lambda_1$  and  $\lambda_2$  are the decay constants of the mother and daughter  
 386 nuclei, respectively. Inserting Eq. 4 and Eq. 5 into Eq. 3, and rearranging  
 387 terms:

$$P_n = \frac{N_{\beta n} - B_{\beta n} - C}{\epsilon_n N_\beta - C}. \quad (6)$$

388 The value of  $N_\beta$  for a given nucleus is calculated as the product of the total  
 389 number of implantations in the DSSD, and the  $\beta$ -detection efficiency.

390 The NERO background rate given in Sec. 3.3 is the “free” neutron back-  
 391 ground rate without any coincidence requirements. In practice, however,  $P_n$



392 values are determined from neutrons measured in coincidence with  $\beta$  decays.  
 393 The number of background  $\beta$ -neutron coincidences for a given nucleus can be  
 394 written as  $B_{\beta n} = B_n(\beta) + B_n(B_\beta)$ ; where  $B_n(\beta)$  is the number of “free” back-  
 395 ground neutrons in random coincidence with parent  $\beta$  decays, and  $B_n(B_\beta)$   
 396 is the number of “free” background neutrons in random coincidence with  
 397 background  $\beta$ -like events in the BCS.

398  $B_n(B_\beta)$  as a function of time and detector pixel can be reliably estimated  
 399 from the  $\beta$ -neutron coincidence rates outside of the correlation window of  
 400 any ion implantations.  $B_n(B_\beta)$  for a specific parent nuclide can then be  
 401 calculated by summing the specific backgrounds at the time and location of  
 402 each individual ion implantation event. The high granularity of the DSSD  
 403 detector greatly reduces this background.

404  $B_n(\beta)$  can be calculated as the product of the number of parent  $\beta$  decays  
 405 detected  $N_\beta$  and the probability for at least one “free” background neutron  
 406 to be detected in random coincidence with each parent  $\beta$  decay, i.e.:

$$B_n(\beta) = N_\beta \sum_{k=1}^{\infty} \frac{(R_b \tau)^k}{k!} e^{-R_b \tau} = N_\beta (1 - e^{-R_b \tau}), \quad (7)$$

407 where  $\tau$  is the TDC time window defined in Sec. 3.1, and  $R_b$  is the “free”  
 408 neutron background rate. For a typical value  $R_b \simeq 10 \text{ s}^{-1}$ , about 0.2% of  
 409 the detected parent  $\beta$  decays are in random coincidence with a background  
 410 neutron, setting the order of magnitude of the lowest  $P_n$  values that can be  
 411 measured with NERO under these conditions.

412 The  $P_n$  values and their errors obtained in various NSCL experiments can  
 413 be found in Refs. [5, 6]. The  $P_n$  values measured in these experiments agree  
 414 well with perviously measured well established data.

## 415 5. Summary and conclusions

416 The neutron detector NERO has been built at NSCL enabling the mea-  
417 surement of  $\beta$ -delayed neutron emission probabilities of r-process nuclei with  
418 fast rare isotope beams. The specific design was motivated by the require-  
419 ment of achieving a high, energy-independent neutron detection efficiency  
420 up to  $\approx 1$  MeV, and accommodating a large pixelated  $\beta$ -counting system,  
421 necessary to perform measurements with fragmentation beams. MCNP sim-  
422 ulations were carried out during the design phase to find the optimum con-  
423 figuration.

424 Studies of the detector efficiency at various neutron energies were per-  
425 formed with a  $^{252}\text{Cf}$  source, and with neutrons from a number of resonant  
426 and non-resonant reactions at ISNAP. MCNP calculations reproduce rea-  
427 sonably well the energy dependence of the detector efficiency. On the other  
428 hand, the MCNP calculations slightly overestimate the absolute efficiency of  
429 the second and third rings, when the neutrons traverse a larger volume of  
430 polyethylene. An overall scaling of the calculated efficiency to the measured  
431 data can be used to extrapolate the detector efficiency to smaller and larger  
432 neutron energies. The small energy dependence (of about 5%), for neutron  
433 energies below 800 keV, represents the main contribution from the efficiency  
434 correction to the total uncertainty of  $P_n$ .

435 NERO is currently used with the BCS at NSCL, but can be used with  
436 other  $\beta$ -decay stations at other rare isotope facilities. It will also be used to  
437 fully exploit the much higher production rates expected in new generation  
438 facilities like FRIB at NSCL, FAIR at GSI and RIBF at RIKEN.

439 **Acknowledgments**

440 This work was supported in part by the Joint Institute for Nuclear As-  
441 trophysics (JINA) under NSF Grant PHY-02-16783 and the National Super-  
442 conducting Cyclotron Laboratory (NSCL) under NSF Grant PHY-01-10253.

443 **References**

- 444 [1] K.-L. Kratz, J.-P. Bitouzet, F.-K. Thielemann, P. Möller, and B. Pfeiffer,  
445 Ap. J. 403 (1993) 216.
- 446 [2] E.M. Burbidge, G.R. Burbidge, W.A. Fowler, and F. Hoyle, Rev. Mod.  
447 Phys. 29 (1957) 547.
- 448 [3] A.G.W. Cameron, Publ. Astron. Soc. Pac. 69 (1957) 201.
- 449 [4] F. Farouqi, K.-L. Kratz, B. Pfeiffer, T. Rauscher, and F.-K. Thielemann,  
450 AIP Conf. Proc. 819 (2006) 419.
- 451 [5] F. Montes, A. Estrade, P.T. Hosmer, S.N. Liddick, P.F. Mantica,  
452 A.C. Morton, W.F. Mueller, M. Ouellette, E. Pellegrini, P. Santi,  
453 H. Schatz, A. Stolz, B.E. Tomlin, O. Arndt, K.-L.Kratz, B. Pfeiffer,  
454 P. Reeder, W.B. Walters, A. Aprahamian, and A. Wöhr, Phys. Rev. C  
455 73 (2006) 035801.
- 456 [6] J. Pereira, S. Hennrich, A. Aprahamian, O. Arndt, A. Becerril, T. Elliot,  
457 A. Estrade, D. Galaviz, R. Kessler, K.-L.Kratz, G. Lorusso, P.F. Man-  
458 tica, M. Matos, P. Möller, F. Montes, B. Pfeiffer, H. Schatz, F. Schertz,  
459 L. Schnorrenberger, E. Smith, A. Stolz, M. Quinn, W.B. Walters, and  
460 A. Wöhr, Phys. Rev. C 79 (2009) 035806.

- 461 [7] G. Grosshög, in: D.N. Poenaru, W. Greiner (Eds.), Experimental Tech-  
462 niques in Nuclear Physics, Walter de Gruyter, 1997, Chap. 7, p. 235.
- 463 [8] G. Rudstam, K. Aleklett, and L. Sihver, At. Data Nucl. Data Tables 53  
464 (1993) 1.
- 465 [9] . B. Pfeiffer, K.-L. Kratz, and P. Möller, Prog. in Nucl. En. 41 (2002)  
466 39
- 467 [10] T. Mehren, B. Pfeiffer, S. Schoedder, K.-L. Kratz, M. Huhta, P. Den-  
468 dooven, A. Honkanen, G. Lhersonneau, M. Oinonen, J.-M. Parmonen,  
469 H. Penttilä, A. Popov, V. Rubchenya, and J. Äystö, Phys. Rev. Lett.  
470 77 (1996) 458.
- 471 [11] Computer code MCNP 5.0, ORNL (RCICC), Oak Ridge, TN, 2003.
- 472 [12] J.S. Hendricks, S.C. Frankle, and J.D. Court, ENDF/B-VI data for  
473 MCNP, LANL report LA-12891, Los Alamos, 1994.
- 474 [13] J. Elson (private communication); <http://pico-systems.com>.
- 475 [14] O.I. Batenkov, A.B. Blinov, M.V. Blinov, and S.N. Smirnov, Atomic  
476 Energy 64 (2005) 489.
- 477 [15] J.K. Bair and F.X. Haas, Phys. Rev. C 7 (1973) 1356.
- 478 [16] E. Ramström and T. Wiedling, Nucl. Phys. A 272 (1976) 259.
- 479 [17] C.R. Brune, I. Licot, and R.W. Kavanagh, Phys. Rev. C 48 (1993) 3119.
- 480 [18] E. Lund, P. Hoff, K. Aleklett, O. Glomset, and G. Rudstam, Z. Phys.  
481 A 294 (1980) 233.

- 482 [19] T.R. Wang, R.B. Vogelaar, and R.W Kavanagh, Phys. Rev. C 43 (1991)  
483 883.
- 484 [20] W. A. Fowler, C. C. Lauritsen, and T. Lauritsen, Rev. Mod. Phys. 20  
485 (1948) 236.
- 486 [21] J.F. Ziegler, J.P. Biersack, and M.D. Ziegler, The Stopping and Ranges  
487 of Ions in Matter, Lulu Press, Morrisville NC, 2008.
- 488 [22] J.L. Zyskind C.A. Barnes, J.M. Davidson, W.A. Fowler, R.E. Marrs,  
489 and M.H. Shapiro, Nucl. Phys. A 343 (1980) 295.
- 490 [23] P. Yalçın and Y. Kurucu, Applied Radiation and Isotopes 62 (2005) 63.
- 491 [24] R.G. Helmer and C. van der Leun, Nucl. Instrum. and Meth. in Phys.  
492 Res. A 450 (2000) 35.
- 493 [25] C. Iliadis, *Nuclear Physics of Stars* (Wiley-VCH, Weinheim, 2007) Ap-  
494 pendix D.2.
- 495 [26] J.P. Schiffer, A.A. Kraus, and J.R. Risser, Phys. Rev. 105 (1957) 1811.
- 496 [27] W. Bröchle, Radiochimica Acta 91 (2003) 71.
- 497 [28] Y. Danon, R.C. Block, and R.E. Slovacek, Nucl. Instrum. Methods Phys.  
498 Res. A 352 (1995) 596.
- 499 [29] ESARDA NDA Working Group, ESARDA Bulletin 33 (November 2003)  
500 26.
- 501 [30] B. Wolle, R. Bätzner, T. Baloui, G. Gonda, H. Klein, B. Wiegel, and  
502 J. Wittstock, Rev. Sci. Instrum. 70 (1999) 1194.

- 503 [31] K.-L. Kratz, in Proceedings of the consultants' meeting on delayed neu-  
504 tron properties (IAEA, Vienna, 1979) p. 103
- 505 [32] K.-L. Kratz, W. Rudolph, H. Ohm, H. Franz, M. Zendel, G. Hermann,  
506 S.G. Prussin, F.M. Nuh, A.A. Shihab-Eldin, D.R. Slaughter, W. Halver-  
507 son, and H.V. Klapdor, Nucl. Phys. A 317 (1979) 335.
- 508 [33] K.-L. Kratz, A. Schröder, H. Ohm, M. Zendel, H. Gabelmann,  
509 W. Ziegert, P. Peuser, G. Jung, B. Pfeiffer, K.D. Wünsch, H. Wollnik,  
510 C. Ristori, and J. Crançon, Z. Phys. A 306 (1982) 239.
- 511 [34] T.R. England, E.D. Arthur, M.C. Brady, and R.J. LaBauve, LA-11151-  
512 MS (1998).
- 513 [35] P.T. Hosmer, H. Schatz, A. Aprahamian, O. Arndt, R.R.C. Clement,  
514 A. Estrade, K.-L.Kratz, S.N. Liddick, P.F. Mantica, W.F. Mueller,  
515 F. Montes, A.C. Morton, M. Ouellette, E. Pellegrini, B. Pfeiffer,  
516 P. Reeder, P. Santi, M. Steiner, A. Stolz, B.E. Tomlin, W.B. Walters,  
517 and A. Wöhr, Phys. Rev. Lett. 94 (2005) 112501.
- 518 [36] P.T. Hosmer, H. Schatz *et al.* (in preparation).
- 519 [37] M. Quinn, J. Pereira *et al.* (in preparation).
- 520 [38] J. Cetnar, Ann. of Nucl. En. 33 (2006) 640.